

# QUANTUM UNIVERSE

---

THE REVOLUTION IN 21<sup>ST</sup> CENTURY PARTICLE PHYSICS

DOE / NSF

HIGH ENERGY PHYSICS ADVISORY PANEL

QUANTUM UNIVERSE COMMITTEE

# QUANTUM UNIVERSE

THE REVOLUTION IN 21ST CENTURY PARTICLE PHYSICS

**What does “Quantum Universe” mean?**

To discover what the universe is made of and how it works is the challenge of particle physics. Quantum Universe presents the quest to explain the universe in terms of quantum physics, which governs the behavior of the microscopic, subatomic world. It describes a revolution in particle physics and a quantum leap in our understanding of the mystery and beauty of the universe.

COMMITTEE MEMBERS

ANDREAS ALBRECHT

University of California at Davis

SAMUEL ARONSON

Brookhaven National Laboratory

KEITH BAKER

Hampton University

Thomas Jefferson National Accelerator Facility

JONATHAN BAGGER

Johns Hopkins University

NEIL CALDER

Stanford Linear Accelerator Center

Stanford University

PERSIS DRELL, CHAIR

Stanford Linear Accelerator Center

Stanford University

EVALYN GATES

University of Chicago

FRED GILMAN

Carnegie Mellon University

JUDITH JACKSON

Fermilab

STEVEN KAHN

Physics Department, Stanford University

Stanford Linear Accelerator Center

EDWARD KOLB

Fermilab

University of Chicago

JOSEPH LYKKEN

Fermilab

HITOSHI MURAYAMA

Institute for Advanced Study, Princeton

University of California, Berkeley

HAMISH ROBERTSON

University of Washington

JAMES SIEGRIST

Lawrence Berkeley National Laboratory

University of California, Berkeley

SIMON SWORDY

University of Chicago

JOHN WOMERSLEY

Fermilab

CONTENTS

iii

EXECUTIVE SUMMARY

1

I

INTRODUCTION

2

II

THE FUNDAMENTAL NATURE OF  
MATTER, ENERGY, SPACE AND TIME

4

EINSTEIN'S DREAM OF UNIFIED FORCES

10

THE PARTICLE WORLD

15

THE BIRTH OF THE UNIVERSE

20

III

TOOLS FOR A SCIENTIFIC REVOLUTION

40

IV

CONCLUSION

42

V

SUMMARY TABLES

44

CHARGE TO THE COMMITTEE

46

FACILITIES GLOSSARY



## EXECUTIVE SUMMARY

### WHAT IS THE NATURE OF THE UNIVERSE AND WHAT IS IT MADE OF?

### WHAT ARE MATTER, ENERGY, SPACE AND TIME?

### HOW DID WE GET HERE AND WHERE ARE WE GOING?

Throughout human history, scientific theories and experiments of increasing power and sophistication have addressed these basic questions about the universe. The resulting knowledge has led to revolutionary insights into the nature of the world around us.

In the last 30 years, physicists have achieved a profound understanding of the fundamental particles and the physical laws that govern matter, energy, space and time. Researchers have subjected this “Standard Model” to countless experimental tests; and, again and again, its predictions have held true. The series of experimental and theoretical breakthroughs that combined to produce the Standard Model can truly be celebrated as one of the great scientific triumphs of the 20th century.

Now, in a development that some have compared to Copernicus’s recognition that the earth is not the center of the solar system, startling new data have revealed that only five percent of the universe is made of normal, visible matter described by the Standard Model. Ninety-five percent of the universe consists of dark matter and dark energy whose fundamental nature is a mystery. The Standard Model’s orderly and elegant view of the universe must be incorporated into a deeper theory that can explain the new phenomena. The result will be a revolution in particle physics as dramatic as any that have come before.

### QUESTIONS FOR THE UNIVERSE

A worldwide program of particle physics investigation is underway to explore the mysterious new scientific landscape. Nine interrelated questions define the path ahead.

EINSTEIN'S DREAM OF UNIFIED FORCES

1

ARE THERE UNDISCOVERED PRINCIPLES OF NATURE :  
NEW SYMMETRIES, NEW PHYSICAL LAWS?

The quantum ideas that so successfully describe familiar matter fail when applied to cosmic physics. Solving the problem requires the appearance of new forces and new particles signaling the discovery of new symmetries–undiscovered principles of nature’s behavior.

2

HOW CAN WE SOLVE THE MYSTERY OF DARK ENERGY?

The dark energy that permeates empty space and accelerates the expansion of the universe must have a quantum explanation. Dark energy might be related to the Higgs field, a force that fills space and gives particles mass.

3

ARE THERE EXTRA DIMENSIONS OF SPACE?

String theory predicts seven undiscovered dimensions of space that give rise to much of the apparent complexity of particle physics. The discovery of extra dimensions would be an epochal event in human history; it would change our understanding of the birth and evolution of the universe. String theory could reshape our concept of gravity.

4

DO ALL THE FORCES BECOME ONE?

At the most fundamental level all forces and particles in the universe may be related, and all the forces might be manifestations of a single grand unified force, realizing Einstein’s dream.

THE PARTICLE WORLD

5

WHY ARE THERE SO MANY KINDS OF PARTICLES?

Why do three families of particles exist, and why do their masses differ so dramatically? Patterns and variations in the families of elementary particles suggest undiscovered underlying principles that tie together the quarks and leptons of the Standard Model.

6

WHAT IS DARK MATTER?

HOW CAN WE MAKE IT IN THE LABORATORY?

Most of the matter in the universe is unknown dark matter, probably heavy particles produced in the big bang. While most of these particles annihilated into pure energy, some remained. These remaining particles should have a small enough mass to be produced and studied at accelerators.

7

WHAT ARE NEUTRINOS TELLING US?

Of all the known particles, neutrinos are the most mysterious. They played an essential role in the evolution of the universe, and their tiny nonzero mass may signal new physics at very high energies.

THE BIRTH OF THE UNIVERSE

8

HOW DID THE UNIVERSE COME TO BE?

According to cosmic theory, the universe began with a singular explosion followed by a burst of inflationary expansion. Following inflation, the universe cooled, passing through a series of phase transitions and allowing the formation of stars, galaxies and life on earth. Understanding inflation requires breakthroughs in quantum physics and quantum gravity.

9

WHAT HAPPENED TO THE ANTIMATTER?

The big bang almost certainly produced equal amounts of matter and antimatter, yet the universe seems to contain no antimatter. How did the asymmetry arise?

OPPORTUNITIES FOR DISCOVERY

We live in an age when the exploration of great questions is leading toward a revolutionary new understanding of the universe.

“Opportunities have emerged for discovery about the fundamental nature of the universe that we never expected,” Presidential Science Advisor John Marburger said recently. “Technology places these discoveries within our reach, but we need to focus efforts across widely separated disciplines to realize the new opportunities.”

*Quantum Universe* is a response to that challenge. It serves as a guide to where the search for understanding has taken us so far, and to where it is going. The chapters that follow articulate how existing and planned particle physics experiments at accelerators and underground laboratories, together with space probes and ground-based telescopes, bring within reach new opportunities for discovery about the fundamental nature of the universe.

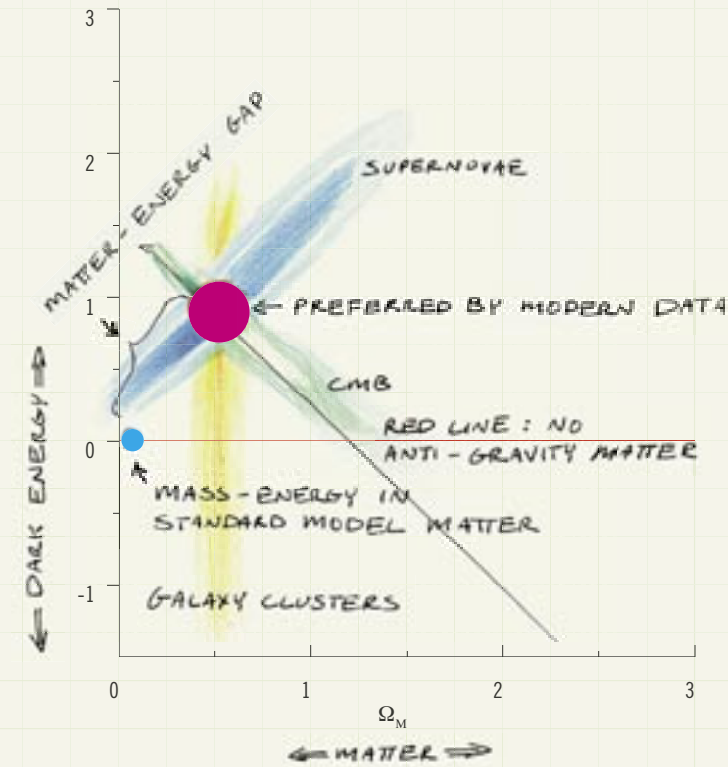
DARK MATTER AND DARK ENERGY

What is the universe made of? The graph at right shows data from several sources that have ruled out long-held ideas about how much and what kinds of matter and energy the universe contains. The data tell us that the total matter-energy content of the universe must include invisible dark matter that holds the universe together, and a mysterious dark energy that pushes the universe apart. The vertical axis shows the amount of dark energy. The horizontal axis shows the amount of matter, both visible and dark.

**Clusters of Galaxies**  
Clusters of galaxies are the most massive objects in the universe. Beside galaxies, these systems contain large quantities of hot gas, which emits x-rays. The temperature of the x-ray-emitting gas, and the speeds of the orbiting galaxies let us infer a cluster's total mass, which is much larger than the mass of the galaxies or the hot gas. This extra mass is in the form of dark matter; and measurements of clusters tell us about the total dark matter content of the universe. The abundance of clusters as a function of cosmic time also gives us information on

**CMB**  
The Cosmic Microwave Background (CMB) is light emitted from the hot plasma of the big bang almost 14 billion years ago—an “echo” of the big bang. To reach us, this image of the ancient universe must propagate for billions of years as the universe expands. Measurable patterns in the CMB teach us about the matter and energy content of the universe as it was at the time when the light was emitted and about how it has changed over time. Models of cosmic matter and energy outside the green concentric ellipses do not fit with data from the CMB.

how much dark energy is present. Models of matter and energy outside the orange ellipses cannot account for the observed properties of clusters.  
  
**Supernovae**  
Supernovae, exploding stars that briefly shine as brightly as galaxies, are observable at such great distances that they can be used to measure the evolution of the universe. One type of supernova appears always to emit the same amount of light at its peak brightness. By measuring the brightness of one of these supernovae, we can tell how far away it is. Comparing this distance with the observed velocity of the supernova tells us how the universe is expanding with



cosmic time. Models of mass and energy outside the blue ellipses predict evolution of the universe that is inconsistent with supernovae observations.

**Standard Model Matter**  
Laboratory experiments have pinned down the behavior of known matter particles. With this knowledge, we can determine how particle interactions affect important processes in the universe. The amounts of light elements like deuterium, helium, and lithium made in the big bang and the propagation of sound wave in the early universe, observable as patterns in the CMB, precisely determine the total amount

of “standard model matter” that the universe contains. The dark blue dot shows the contribution of this matter to the total matter and energy of the universe.

**What’s It Made Of?**  
The pink dot shows the total matter and energy required to agree with all the modern cosmological data. Observations from many sources exclude a universe made only of ordinary matter. They require dark matter and a large amount of dark energy to fill the “matter-energy gap” and achieve a matter and energy total that agrees with the data.

CHAPTER I  
INTRODUCTION

WHAT IS THE NATURE OF THE UNIVERSE  
AND WHAT IS IT MADE OF?

WHAT ARE MATTER, ENERGY,  
SPACE AND TIME?

HOW DID WE GET HERE AND  
WHERE ARE WE GOING?

The quest to answer the most basic questions about the universe has reached a singular moment. As the 21st century begins, physicists have developed a commanding knowledge of the particles and forces that characterize the ordinary matter around us. At the same time, astrophysical and cosmological space observations have revealed that this picture of the universe is incomplete—that 95 percent of the cosmos is not made of ordinary matter, but of a mysterious something else: dark matter and dark energy. We have learned that in fact we do not know what most of the universe is made of.

Understanding this unknown “new” universe requires the discovery of the particle physics that determines its fundamental nature. Powerful tools exist to bring the physics within reach. With astrophysical observations, we can explore the parameters of the universe; with accelerator experiments we can search for their quantum explanation. Energies at particle accelerators now approach the conditions in the first instants after the big bang, giving us the means to discover what dark matter and dark energy are—and creating a revolution in our understanding of particle physics and the universe.

To answer the fundamental questions about the nature of the universe, astrophysical observations of the relics of the big bang must agree with data from physics experiments recreating the particles and forces of the early universe. The two ends of the exploration must meet. We will answer these challenging questions by combining what we learn from the most powerful and insightful observations and experiments in each of these approaches.

The following chapters take up in successively greater depth nine key questions that define the field. Chapter II elucidates the meaning of the questions and their implications for 21<sup>st</sup> century particle physics; Chapter III defines the scientific program to address and answer them. The tables in Chapter V summarize the U.S. facilities whose primary physics programs respond most directly to the questions.





## QUESTIONS FOR THE UNIVERSE

### EINSTEIN'S DREAM OF UNIFIED FORCES

1

ARE THERE UNDISCOVERED PRINCIPLES OF NATURE :  
NEW SYMMETRIES, NEW PHYSICAL LAWS?

2

HOW CAN WE SOLVE THE MYSTERY OF DARK ENERGY?

3

ARE THERE EXTRA DIMENSIONS OF SPACE?

4

DO ALL THE FORCES BECOME ONE?

### THE PARTICLE WORLD

5

WHY ARE THERE SO MANY KINDS OF PARTICLES?

6

WHAT IS DARK MATTER?

HOW CAN WE MAKE IT IN THE LABORATORY?

7

WHAT ARE NEUTRINOS TELLING US?

### THE BIRTH OF THE UNIVERSE

8

HOW DID THE UNIVERSE COME TO BE?

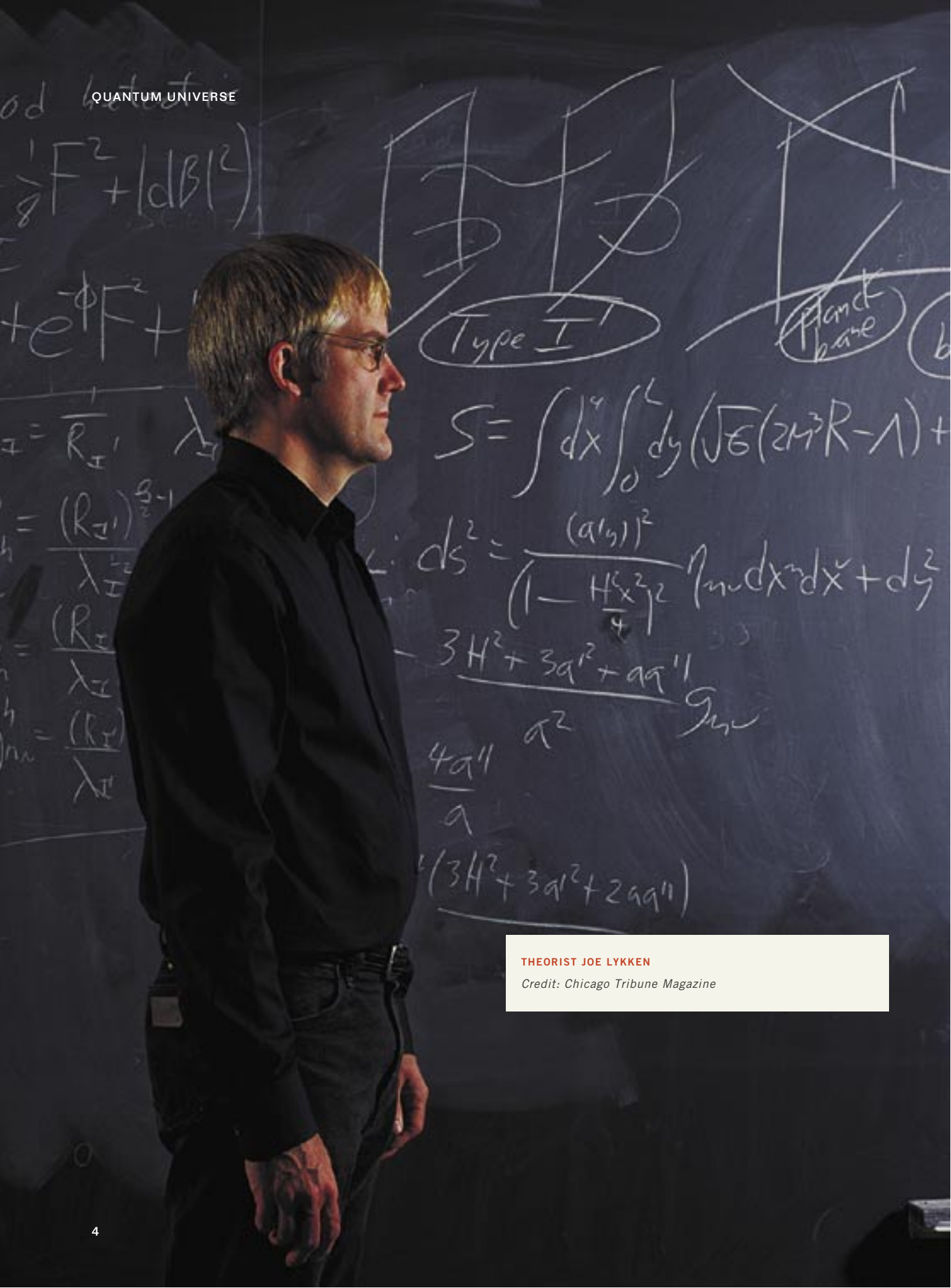
9

WHAT HAPPENED TO THE ANTIMATTER?

*Photo Credit: Allan Morton/Dennis Millon/SPL*

## CHAPTER II THE FUNDAMENTAL NATURE OF MATTER, ENERGY, SPACE AND TIME

Current and future particle physics experiments around the world give us the capability to address a well-defined set of questions about the basic physical laws that govern the universe. These questions, at the same time eminently familiar and profoundly revolutionary, define the path for particle physics in the 21<sup>st</sup> century.



QUANTUM UNIVERSE

THEORIST JOE LYKKEN

Credit: Chicago Tribune Magazine

- 1  
ARE THERE UNDISCOVERED PRINCIPLES OF NATURE :  
NEW SYMMETRIES, NEW PHYSICAL LAWS?
- 2  
HOW CAN WE SOLVE THE MYSTERY OF DARK ENERGY?
- 3  
ARE THERE EXTRA DIMENSIONS OF SPACE?
- 4  
DO ALL THE FORCES BECOME ONE?

## EINSTEIN'S DREAM OF UNIFIED FORCES

Since Einstein, physicists have sought a unified theory to explain all the fundamental forces and particles in the universe. The result is a stunningly successful theory that reduces the complexity of microscopic physics to a set of concise laws. But these same quantum ideas fail when applied to cosmic physics. Some fundamental piece is missing; gravity, dark matter and dark energy must have quantum explanations. A new theoretical vision is required, one that embraces the Standard Model and general relativity, while resolving the mystery of dark energy. Particle accelerators provide the means to reach a unified theoretical perspective in experiments characterized by four well-defined intellectual thrusts.



# 1

EINSTEIN'S DREAM OF UNIFIED FORCES

## ARE THERE UNDISCOVERED PRINCIPLES OF NATURE : NEW SYMMETRIES, NEW PHYSICAL LAWS?

Our quest to discover the fundamental laws of nature has led to the revelation that the laws of physics, and the particles they govern, exist because of underlying symmetries of nature, some of them lost since the big bang. One such lost symmetry might be supersymmetry. Just as for every particle there exists an antiparticle, supersymmetry predicts that for every known particle there also exists a superpartner particle. Part of the strong theoretical appeal of supersymmetry, an essential part of string theory, is its possible connection to dark energy and the fact that it provides a natural candidate for dark matter, the neutralino.

The discovery of supersymmetry is an immediate experimental challenge of particle physics, followed by the exploration of its structure and the properties of the superpartner particles. Particle accelerator experiments will uncover the role of supersymmetry in a unified theory and reveal whether the neutralino superpartner accounts for dark matter.

# 2

EINSTEIN'S DREAM OF UNIFIED FORCES

## HOW CAN WE SOLVE THE MYSTERY OF DARK ENERGY?

Recent measurements with telescopes and space probes have shown that a mysterious force—a dark energy—fills the vacuum of empty space, accelerating the universe’s expansion. We don’t know what dark energy is, or why it exists. On the other hand, particle theory tells us that, at the microscopic level, even a perfect vacuum bubbles with quantum particles that are a natural source of dark energy. But a naïve calculation of the dark energy generated from the vacuum yields a value  $10^{120}$  times larger than the amount we observe. Some unknown physical process is required to eliminate most, but not all, of the vacuum energy, leaving enough left to drive the accelerating expansion of the universe. A new theory of particle physics is required to explain this physical process.

Particle physics data point to another mysterious component of empty space, the Higgs field, that gives particles the property of mass. Without the Higgs field, electrons would fly at the speed of light, and atoms would instantly disintegrate. Are dark energy and the Higgs field related? The discovery of supersymmetry would provide crucial evidence of a possible connection. Supersymmetry provides both a natural context for the Higgs field and a possible explanation for the small but finite value of dark energy.

# 3

EINSTEIN'S DREAM OF UNIFIED FORCES

## ARE THERE EXTRA DIMENSIONS OF SPACE?

The revolutionary concept of string theory is a bold realization of Einstein's dream of an ultimate explanation for everything from the tiniest quanta of particle physics to the cosmos itself. String theory unifies physics by producing all known forces and particles as different vibrations of a single substance called superstrings. String theory brings quantum consistency to physics with an elegant mathematical construct that appears to be unique.

Do superstrings exist? The strings themselves are probably too tiny to observe directly, but string theory makes a number of testable predictions. It implies supersymmetry and predicts seven undiscovered dimensions of space, dimensions that would give rise to much of the mysterious complexity of particle physics. Testing the validity of string theory requires searching for the extra dimensions and exploring their properties. How many are there? What are their shapes and sizes? How and why are they hidden? And what are the new particles associated with the extra dimensions?

# 4

EINSTEIN'S DREAM OF UNIFIED FORCES

## DO ALL THE FORCES BECOME ONE?

At the most fundamental level, particles and forces may converge, either through hidden principles like grand unification, or through radical physics like superstrings. We already know that remarkably similar mathematical laws and principles describe all the known forces except gravity. Perhaps all forces are different manifestations of a single grand unified force, a force that would relate quarks to leptons and predict new ways of converting one kind of particle into another. Such a force might eventually make protons decay, rendering ordinary matter unstable.

5

WHY ARE THERE SO MANY KINDS OF PARTICLES?

6

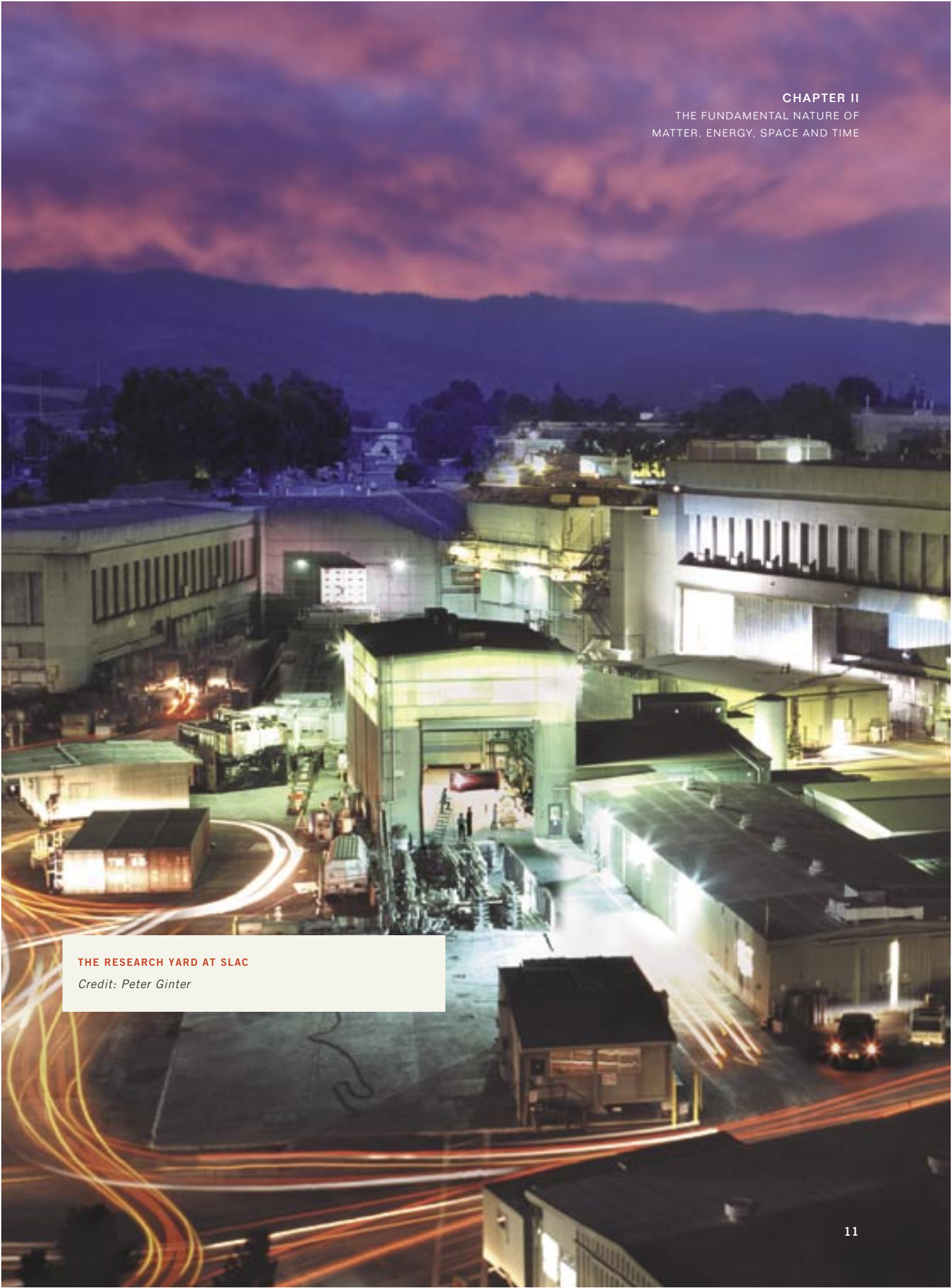
WHAT IS DARK MATTER?  
HOW CAN WE MAKE IT IN THE LABORATORY?

7

WHAT ARE NEUTRINOS TELLING US?

THE PARTICLE WORLD

Physicists have identified 57 distinct species of elementary particles and have determined many of their properties in exquisite detail. What are their roles? How will we know when we have found them all? Perhaps these particles are just different notes on a single superstring. Perhaps they are related by grand unification, or other hidden symmetries, in ways that we have yet to decipher. Unification may provide the key, the simple principle that gives particles their complex identities.



THE RESEARCH YARD AT SLAC  
*Credit: Peter Ginter*



# 5

THE PARTICLE WORLD

## WHY ARE THERE SO MANY KINDS OF PARTICLES?

We have discovered three families of quarks and leptons, families of fundamental particles that differ only in their masses, which range from less than a millionth of the mass of an electron to the mass of an atom of gold. Just as quantum mechanics led to an understanding of the organization of the periodic table, we look to new theories to explain the patterns of elementary particles. Why do three families of particles exist, and why do their masses differ so dramatically?

Current investigations focus on developing a detailed picture of the existing patterns in the particle world. Remarkable progress has been made, especially in characterizing the quarks. But why are the patterns for leptons and quarks completely different? Detailed studies of quarks and leptons at accelerator experiments will provide the clearest insight into these issues.

# 6

THE PARTICLE WORLD

## WHAT IS DARK MATTER? HOW CAN WE MAKE IT IN THE LABORATORY?

Most of the matter in the universe is dark. Without dark matter, galaxies and stars would not have formed and life would not exist. It holds the universe together. What is it?

Although the existence of dark matter was suggested in the 1930s, only in the last 10 to 15 years have scientists made substantial progress in understanding its properties, mostly by establishing what it is not. Recent observations of the effect of dark matter on the structure of the universe have shown that it is unlike any form of matter that we have discovered

or measured in the laboratory. At the same time, new theories have emerged that may tell us what dark matter actually is. The theory of supersymmetry predicts new families of particles interacting very weakly with ordinary matter. The lightest supersymmetric particle could well be the elusive dark matter particle. We need to study dark matter directly by detecting relic dark matter particles in an underground detector and by creating dark matter particles at accelerators, where we can measure their properties and understand how they fit into the cosmic picture.

7

THE PARTICLE WORLD

WHAT ARE NEUTRINOS TELLING US?

Ubiquitous, elusive and full of surprises, neutrinos are the most mysterious of the known particles in the universe. They interact so weakly with other particles that trillions of them pass through our bodies each second without leaving a trace. The sun shines brightly in neutrinos, produced in the internal fusion reactions that power the sun. These reactions produce neutrinos of only one kind, but they mysteriously morph into two other kinds on their way to earth. Neutrinos have mass, but the heaviest neutrino is at least a million times lighter than the lightest charged particle.

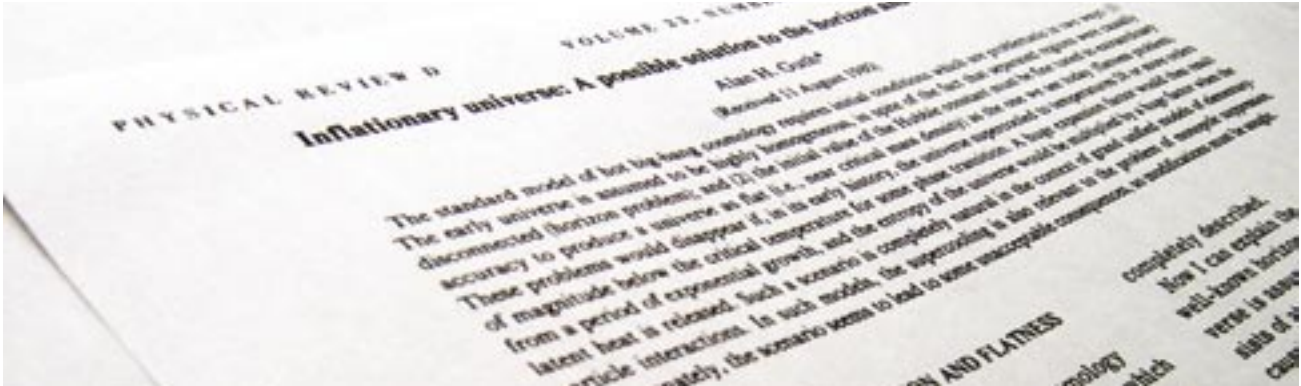
The existence of the neutrino’s tiny nonzero mass raises the possibility that neutrinos get their masses from unknown physics, perhaps related to unification. Detailed studies of the properties of neutrinos—their masses, how they change from one kind to another, and whether neutrinos are their own antiparticles—will tell us whether neutrinos conform to the patterns of ordinary matter or whether they are leading us toward the discovery of new phenomena.

8

HOW DID THE UNIVERSE COME TO BE?

9

WHAT HAPPENED TO THE ANTIMATTER?



THE BIRTH  
OF THE UNIVERSE

What started the big bang? How did space, time, matter and energy take the forms that we see today? Can we work backward to unravel the history of the universe?

After the big bang exploded with enormous energy, the universe began the cool-down that has lasted until our own time. The resulting chain of events is a cosmic drama with many acts, dramatic transitions and a host of actors appearing and disappearing along the way. The early scenes played out at unimaginable temperatures and densities, the stage set by the fundamental properties of particle physics. These processes had to be finely tuned to yield a universe capable of forming the galaxies, stars and planets we observe today. Did some undiscovered fundamental laws determine the conditions that allowed us to exist?

To reconstruct the cosmic story, telescopes and space probes detect the relics from the early universe, and particle accelerators recreate and study the extreme physics that characterized the stages of development and the transitions between them. As we begin to understand the cosmic past, we can look to the future of the universe and predict its ultimate fate.

# 8

THE BIRTH OF THE UNIVERSE

## HOW DID THE UNIVERSE COME TO BE?

According to modern theories of cosmic evolution, the universe began with a singular explosion, followed by a burst of inflationary expansion. To understand inflation requires breakthroughs in our understanding of fundamental physics, of quantum gravity, and of the ultimate unified theory. Although inflationary conditions are too high in energy to reproduce on earth, we can observe their signatures, transmitted over the eons by their imprint on the relic matter we can still detect from that era.

Following inflation, the conditions of the early universe were still so extreme they could combine elementary particles into new phases of matter. As the universe expanded and cooled, transitions took place as matter changed from one phase to another, like steam condensing into water. Some of these phase transitions may have been the most dramatic events in cosmic history, shaping the evolution of the universe and leaving relics observable today. Cosmic phase transitions could be recreated in high-energy accelerator experiments.

# 9

THE BIRTH OF THE UNIVERSE

## WHAT HAPPENED TO THE ANTIMATTER?

Experiments teach us that for every fundamental particle there exists an antiparticle. The big bang and its aftermath almost certainly produced particles and antiparticles in equal numbers. However, for as far out in the universe as we can probe, our observations indicate that we live in a universe of matter, not antimatter. What happened to the antimatter? A tiny imbalance between particles and antiparticles must have developed early in the evolution of the universe, or it all would have annihilated, leaving only photons and neutrinos. Subtle asymmetries between matter and antimatter, some of which we have observed experimentally in the laboratory, must be responsible for this imbalance. But our current knowledge of these asymmetries is incomplete, insufficient to account for the observed matter domination.

There must be some other undiscovered phenomenon that makes matter and antimatter behave differently. We may discover it in quarks—or in neutrinos. Its source may lie in the properties of the Higgs boson, in supersymmetry or even in extra dimensions.



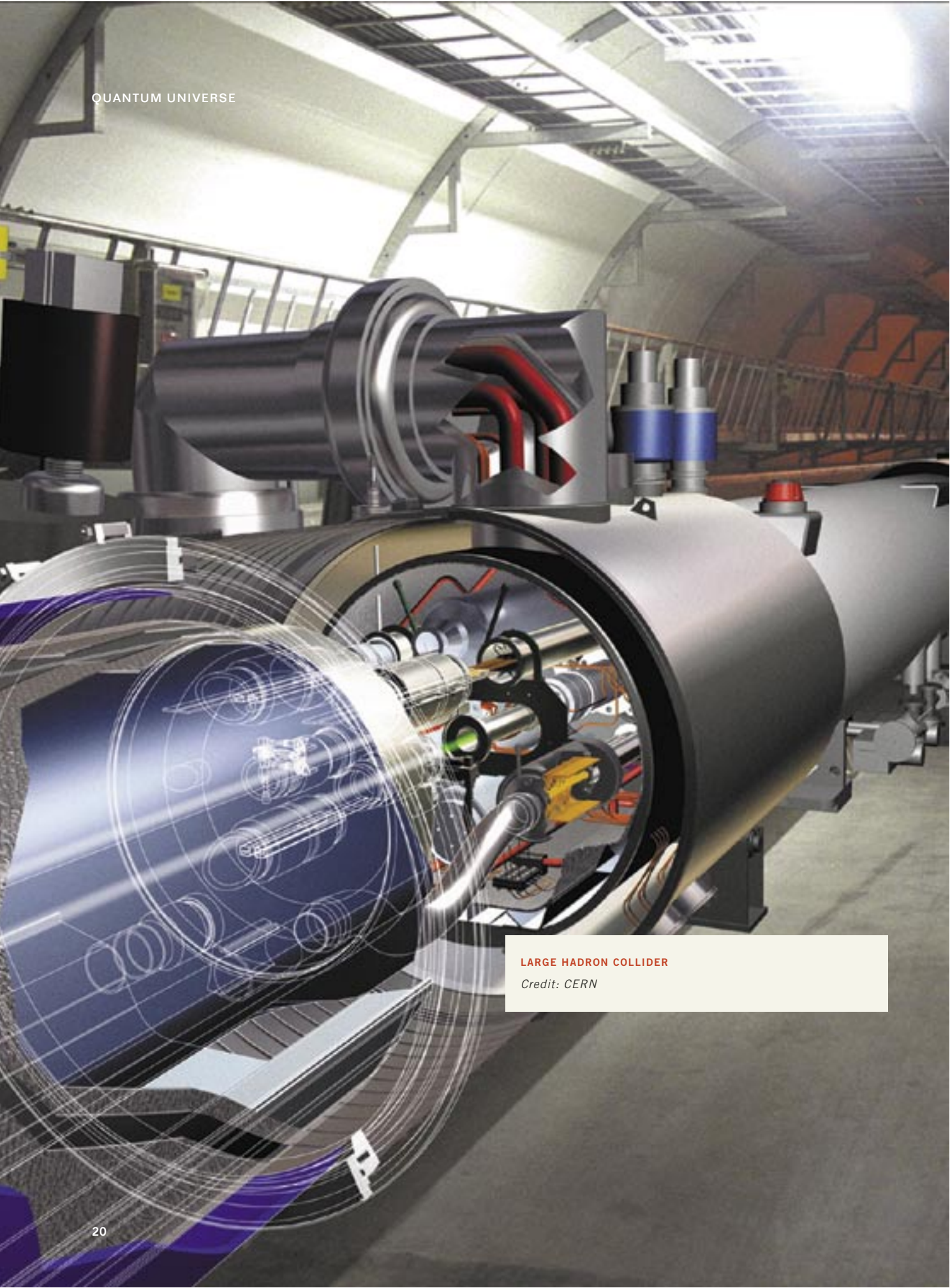
THE GOAL

The nine questions we have asked are the questions that define the science of particle physics today. Answering these questions is the goal of the worldwide program of investigation, driving current and future experiments and the planning for new scientific facilities. In Chapter III we describe the experiments designed to answer these questions.



THEORY THINK TANK  
Credit: Peter Ginter





**LARGE HADRON COLLIDER**  
*Credit: CERN*

# CHAPTER III TOOLS FOR A SCIENTIFIC REVOLUTION

The particle physics community is embarked on an ambitious experimental program that takes the next step toward a revolution in understanding the fundamental nature of matter and energy, space and time. Insights from experiments at particle accelerators, from cosmological observations and from underground experiments have focused the investigation. Further advances require a vigorous program of exploration. Chapter III maps the experimental program onto the questions that define 21<sup>st</sup> century particle physics, articulating the strengths of the different approaches and the scientific goals they are designed to achieve.

At the end of this report, we present two tables that list selected facilities of the U.S. program. The primary physics goals of each of these facilities align most directly with the nine questions of this report. In the text of Chapter III, we also discuss certain other facilities that have the potential to make major contributions to the physics presented here.

EINSTEIN’S DREAM OF UNIFIED FORCES

Superstrings and grand unification are currently the most promising ideas for fulfilling Einstein’s dream of an ultimate theory. To understand string theory requires experimental tests of its predictions of supersymmetric particles and extra dimensions of space. To understand unification requires experiments sensitive to extremely rare particle decays. To confront the discrepancy between the theory of the large and the theory of the small requires a better understanding of fundamental quantum physics.

1  
ARE THERE UNDISCOVERED PRINCIPLES OF NATURE :  
NEW SYMMETRIES, NEW PHYSICAL LAWS?

The laws of nature derive from nature’s symmetries; searching for new particles and forces means searching for new symmetries. One such symmetry might be supersymmetry, which predicts that for every known particle there exists a superpartner particle of the same mass. Experiments have not yet detected any of the superpartners; thus, if supersymmetry exists, it must be broken by unknown physics that makes superpartner particles heavy. Superpartner masses may be related to the Higgs field; supersymmetry provides a natural dark matter candidate, the neutralino. Experiments are searching for supersymmetry now; directly at the Tevatron and indirectly at the B-factories Belle and BaBar.

The Tevatron may have enough energy to produce detectable signals of the lightest superpartners. The LHC should have enough energy to produce all or most of the superpartner particles, either directly or

through the decays of other superpartners, to determine the pattern of superpartner masses and decays.

A Linear Collider would measure the properties of the superpartners very precisely, showing that they are indeed the superpartners of known particles; it could study the properties of the lightest superpartner (most likely the neutralino) with great precision. Do neutralinos behave like dark matter? Studies of the neutralino at a Linear Collider, combined with precision measurements of other superpartners, would produce a prediction for the cosmic relic density of neutralinos to determine whether the predictions are consistent with the dark matter hypothesis.

Theoretical models for the physical mechanism that breaks supersymmetry are already constrained by data from Belle and BaBar. Future precision

studies at Belle and BaBar, as well as from the future hadron B-factories BTeV and LHC-b will allow physicists to disentangle the flavor structure of supersymmetry through subtle changes to decays of B mesons. The MECO experiment will provide unprecedented sensitivity to the direct conversion of muons into electrons in nuclei; and some models of supersymmetric grand unification predict rates for this process that MECO can observe.

2  
HOW CAN WE SOLVE THE MYSTERY OF DARK ENERGY?

The dramatic discovery of dark energy showed that empty space is filled with a mysterious energy that increases as the universe expands. While Einstein initially proposed a cosmological constant that could explain the dark energy, it is the amount of dark energy that is difficult to understand. The natural source of such a dark energy field, quantum fluctuations of the vacuum, gives a density of dark energy  $10^{120}$  times larger than observed levels.

A far-reaching program is in place to study the properties of dark energy. Measurements of the amplitude and fluctuations of the cosmic microwave background from WMAP, combined with data from worldwide astronomical facilities, especially supernova measurements, suggest that dark energy is consistent with a cosmological constant. Future measurements of supernovae, gravitational lensing and clusters of galaxies from a Large Synoptic Survey Telescope, LSST, and the Joint Dark Energy Mission, JDEM, will reveal

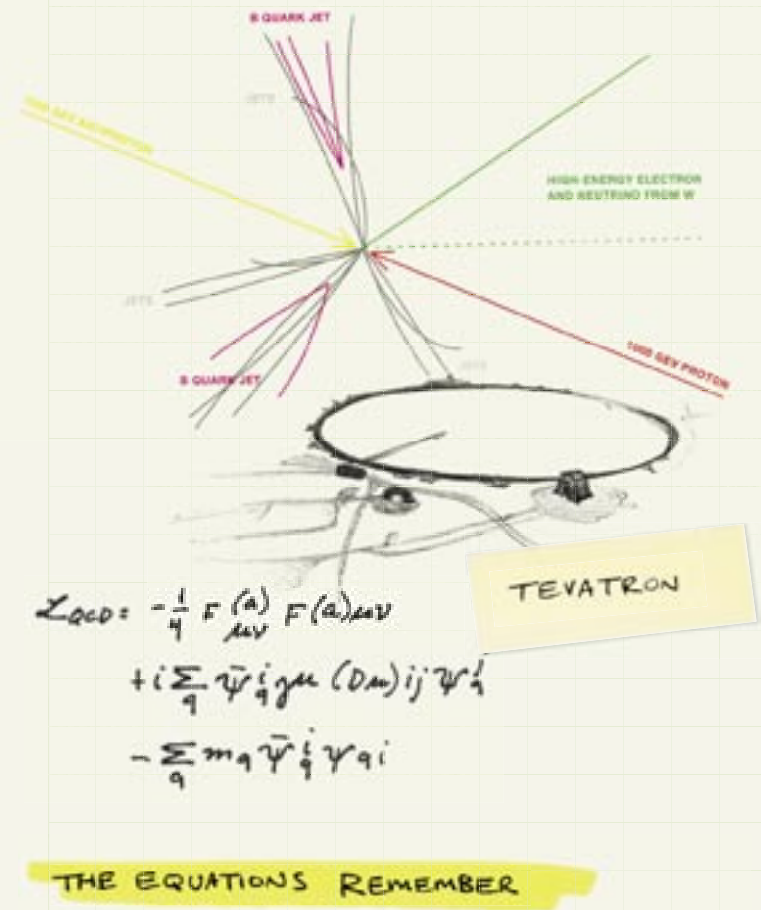


**CMS EXPERIMENT** USCMS is a collaboration of US scientists participating in construction, software and physics analysis at the LHC’s Compact Muon Solenoid experiment. *Credit: CERN*

definitively whether dark energy behaves like Einstein’s cosmological constant or like some new substance that changes with time as the universe evolves.

To determine what dark energy is and why it exists requires connecting the cosmic reality of dark energy to a better fundamental understanding of microscopic quantum physics. At the microscopic scale, physicists have long known that “empty” space is not empty; it is filled by a field that gives quarks and leptons their mass. In the Standard Model, this field is called the Higgs; experiments at the LHC will find the corresponding Higgs particle.





THE EQUATIONS REMEMBER

We live in a cold and empty universe, in a time when energies are so low that we can no longer see what space contained in the fiery instants when the universe was born. But every kind of particle that ever existed is still there, in the equations that describe the particles and forces of the universe. And we can use accelerators to make the equations come alive: not just as metaphor but as reality.

The universe today contains only the stable relics and leftovers of the big bang. The unstable particles have decayed away with time, and the perfect symmetries have been broken as the universe has cooled. But the structure of space “remembers” all the particles and forces we can no longer see around us. Particle accelerators pump energy into empty space to create the particles and uncover the symmetries that existed in the earliest universe. As accelerators go to higher energies, we probe ever closer to the beginning.

Ultimately we must combine what we learn from accelerators with what we learn by detecting the surviving relics of the big bang, using telescopes on the ground and in the sky. It is by synthesizing what we learn from each of these approaches that we will make the two ends meet and develop a comprehensive picture of the universe and how it evolved.

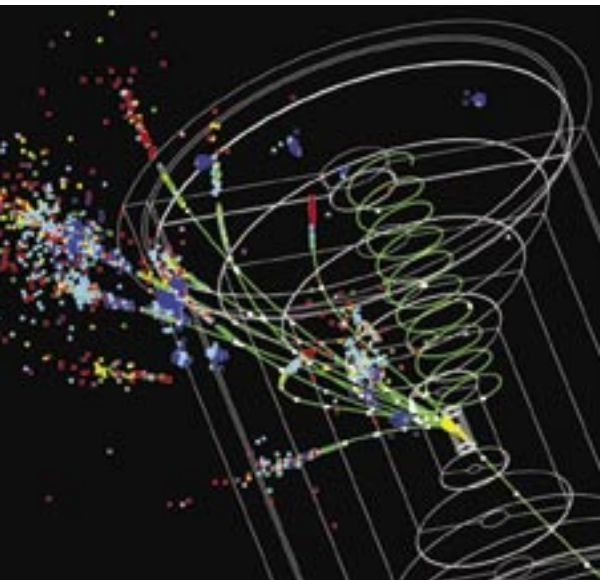
At present, we expect the Higgs to be accompanied by a whole new sector of fundamental physics. This Higgs sector may involve many new particles and interactions. Initial Higgs discoveries will occur at the LHC; a Linear Collider will be essential to explore the landscape of Higgs physics.

Dark energy may have relationships to both supersymmetry and the Higgs sector, implying a new emphasis on the quantum consistency of Higgs physics, including Higgs self-interactions. Such measurements will present additional challenges for the experimental program of the Linear Collider and may provide a foundation to explore the origin of dark energy.

3  
ARE THERE EXTRA DIMENSIONS OF SPACE?

The physical effects of extra dimensions depend on their sizes and shapes, and on what kinds of matter or forces can penetrate them. The sizes of the extra dimensions are unknown, but they should be related to fundamental energy scales of particle physics: the cosmological scale, the density of dark energy, the TeV electroweak scale, or the scale of ultimate unification. It may be possible to infer extra dimensions of macroscopic size from inconsistencies in cosmological observations, or from precision tests of short-range gravitational forces. More likely, the extra dimensions are microscopic, in which case high-energy particle accelerators and cosmic ray experiments are the only ways to detect their physical effects.

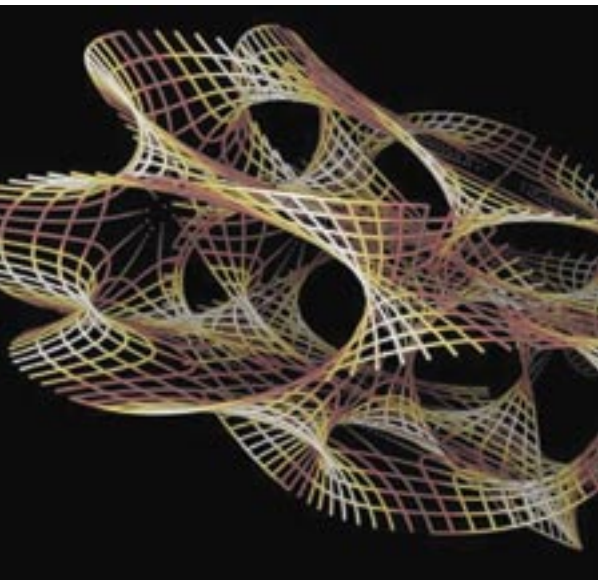
The LHC and a Linear Collider will address many questions about extra dimensions: How many extra dimensions are there? What are their shapes and sizes?



NEUTRALINO Simulation of the signature a neutralino, a predicted supersymmetric particle that is one candidate for dark matter. Credit: Norman Graf

How are they hidden? What are the new particles associated with extra dimensions? Through the production of new particles that move in the extra space, the LHC will have direct sensitivity to extra dimensions 10 billion times smaller than the size of an atom. A Linear Collider would determine the number, size and shape of extra dimensions through their small effects on particle masses and interactions. There is also a chance that, due to the existence of extra dimensions, microscopic black holes may be detected at the LHC or in the highest energy cosmic rays.

Ultimately particle physics seeks to know if dark energy, dark matter and cosmic inflation are affected by the physics of extra dimensions. Collider data will provide insight into the exploration of these deep connections.



SUPERSTRINGS A computer's graphical representation of multi-dimensional spacetime. Credit: Jean-Francois Colonna

4  
DO ALL THE FORCES BECOME ONE?

High-energy particle physics experiments are exploring the unification of the weak and electromagnetic forces. Does the unification continue? At the most fundamental level, particles and forces may be related, either through hidden symmetries like grand unification, or through radical physics like superstrings. Essential clues can come from laboratory observations of extremely rare particle decays and other rare processes, as well as from precision measurements at the highest energies. Such breadth of approach is necessary because we do not know where the clues will appear.



SYMMETRY

The laws of physics distill our observations about what can and cannot happen in the physical world. Powerful conservation laws tell us that certain physical attributes are immune to change. Energy, for example, can take many different forms, but it can be neither created nor destroyed.	An understanding of all the symmetries of nature would be equivalent to knowing how the world works at its most basic level. There are many reasons to believe that we have not found all the symmetries of nature. String theory, for example, requires a new symmetry called supersymmetry. It introduces new quantum dimensions to spacetime and gives meaning to many of the particles we know. It predicts that every elementary particle has a superpartner, just as every particle is accompanied by an antiparticle. Scientists have not yet discovered the supersymmetric partners; finding them is among the greatest challenges for the current and upcoming generations of particle accelerators.
In particle physics, conservation laws result from symmetries in the forces that govern how nature works. A symmetry of the electromagnetic force tells us that electric charge is exactly conserved, that all electrons are created equal, and that the photon—the particle of light—is exactly massless. Such symmetries have provided some of the most dramatic insights in particle physics.	

Models of grand unification predict that protons may eventually decay, rendering ordinary matter unstable. The Super Kamiokande detector already has sensitivity to proton decay in the range suggested by models of grand unification. The next generation of proton decay experiments will require larger detectors in an underground laboratory.

Unification physics represents a natural mechanism to provide Majorana masses for neutrinos, in which case neutrinos are their own antiparticles. This possibility could be verified by a positive signal in the next generation of neutrinoless double beta decay experiments, for example EXO or Majorana, at an underground laboratory.

The precision measurements of the force strengths at LEP and SLC provided circumstantial evidence for grand unification. Similarly the precision measurement of (super)particle masses at a Linear Collider will permit multiple quantitative tests of grand unification.

Ultimately, unification of all the fundamental forces requires an understanding of quantum gravity and the associated exotic phenomena, such as black hole evaporation. Progress could come serendipitously from a discovery of extra dimensions at the LHC, or of anomalous gravity wave sources with LIGO, or of unexpected events in ultra-high-energy cosmic rays, including cosmic neutrinos.

THE PARTICLE WORLD

Physics has revealed that the elementary particles play many roles. The quarks and leptons, the fundamental bits of matter, bind into protons, neutrons and atoms and form the basis of life on earth. The so-called gauge particles give rise to forces, including electricity, magnetism and gravity. Recent discoveries indicate that still more elementary particles make up the dark matter that fills the universe.

In particle physics, elementary particles are the messengers of new phenomena, including new forces and new forms of matter. Quantum mechanics dictates that even hidden dimensions show up as elementary particles. By measuring their properties, physicists can infer the shapes and sizes of the hidden dimensions.

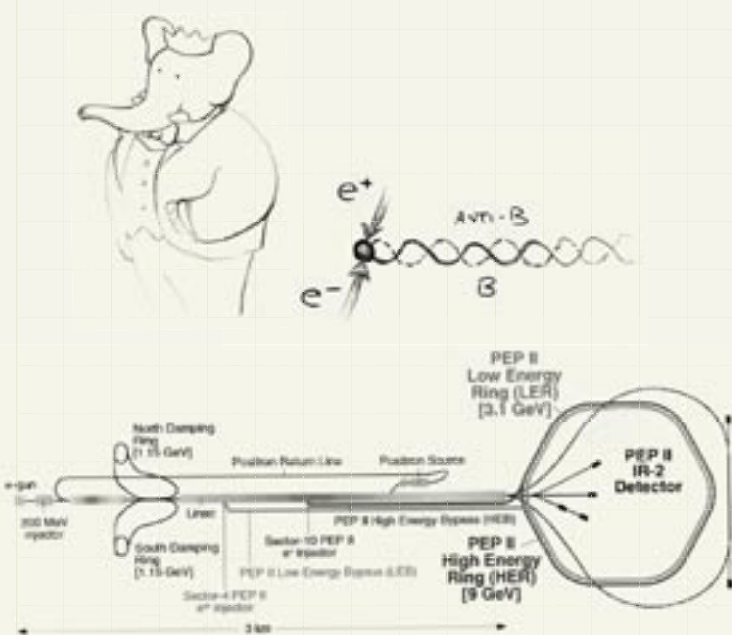
5  
WHY ARE THERE SO MANY KINDS OF PARTICLES?

Physicists have so far identified 57 species of elementary particles. In particular, the Standard Model contains quarks and leptons, grouped into three families that differ only in their masses. Why the pattern of particles is repeated three times with enormous variations in mass but with other properties seemingly identical is an open question. Quantum physics has shown that three families are the minimum necessary to accommodate CP violation in the Standard Model. Such CP violation is necessary for matter to predominate over antimatter in the universe, but its effects observed so far are insufficient to explain this predominance. The current program of experiments focuses on developing a detailed understanding of the existing patterns and searching for signs that the patterns of the three families are not identical.

The CDF and D0 experiments at the Tevatron are measuring the properties of the top quark to see if its enormous mass gives it a special role in the particle world. The BaBar and Belle experiments at SLAC and KEK are using their data samples, containing millions of b and c quarks, as well as  $\tau$  leptons, to make precision measurements of the masses and decay modes of all of these objects, in order to look for subtle deviations from the predicted patterns of their decays. The third-generation particles—top, bottom and tau—offer the best hope for discovery, because their large masses allow them to couple most effectively to undiscovered physics.

BaBar and Belle can study only two types of B mesons, bound states of the bottom quark with up or down quarks. However, many theories suggest significant effects in the bound state with the strange quark,





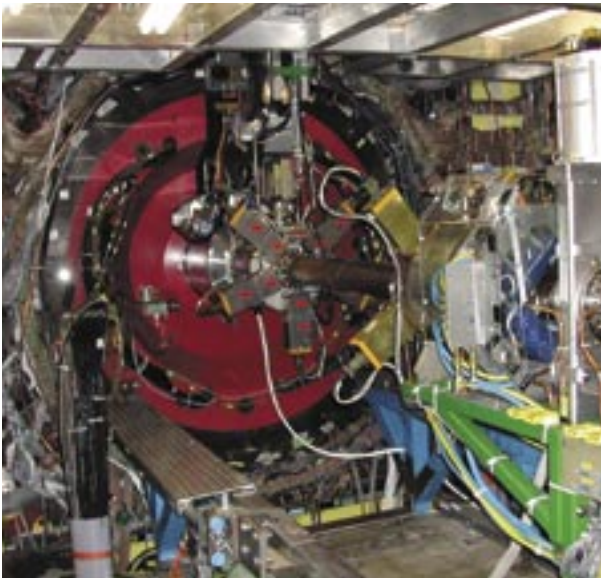
BaBar

Something is missing in physicists’ understanding of how the universe evolved into its current state. At the big bang, equal quantities of matter and antimatter should have been created and subsequently annihilated each other, leaving nothing but energy. However the matter universe is here as undeniable proof of the victory of matter over antimatter in this initial cosmic encounter. To establish experimentally how matter came to dominate is a central theme in particle physics research. The BaBar experiment at SLAC is working relentlessly to tie down an exact measurement of this effect called Charge Parity (CP) violation.

BaBar follows the infinitesimally short lives—a trillionth of a second—of particles called B mesons and those of

their antimatter counterparts, anti-B mesons or “B-bars.” Any difference in behavior of these otherwise exact opposites indicates a difference between matter and antimatter and confirms the existence of CP violation. BaBar’s first results, announced in the summer of 2001, gave clear evidence for CP violation in B mesons. BaBar data continues to provide new levels of understanding of the matter antimatter symmetry.

The success of BaBar is based on the power and efficiency of collaborative science. More than 500 scientists and engineers from 75 institutions in Canada, China, France, Germany, Italy, the Netherlands, Norway, Russia, Spain, the United Kingdom, and the United States are working on BaBar.



**CLEO DETECTOR** CLEO detects, measures and analyzes electron-positron collision events generated by CESR.  
*Credit: Cornell–Wilson Laboratory*

B<sub>s</sub>. Physicists are currently studying the properties of the B<sub>s</sub> meson at the Tevatron. The future hadron B-factories, BTeV and LHC-b, will explore the B<sub>s</sub> meson with far greater precision.

Properties of individual quarks are experimentally difficult to study, because they are always bound to other quarks. Lattice Computational Facilities offer great promise for the calculation of the effects of the strong interactions. As an example, lattice calculations will provide sufficient precision to extract quark parameters, such as those that describe flavor mixing, from the experimental data. Experimental studies with CLEO-c will establish and validate the precision of the lattice calculations for use in heavy quark systems.

Neutrinos have opened a surprising new window on the physics of lepton generations, since neutrino

masses are not necessary in the Standard Model. The presence of neutrino masses may be telling us something about physics beyond the Standard Model. The decay properties of the light leptons, the electron and the muon, may also hold surprises. The MECO experiment proposes to look for a conversion of a muon to an electron and is sensitive to very-high-mass physics that might affect that process.

6  
WHAT IS DARK MATTER?  
HOW CAN WE MAKE IT IN THE LABORATORY?

Most of the matter in the universe is dark. Early evidence for dark matter came from the rotation curves of galaxies, which showed that galaxies contain more mass than is contained in the stars. More recently, direct evidence for dark matter has come from the discovery and characterization of gravitational lenses, regions of space where mass bends light. These astronomical constraints do not directly distinguish between nonbaryonic models for dark matter (WIMPs) and other possible ideas involving more massive objects (MACHOs) such as Jupiter-sized planets and mini-black holes. However experiments in the 1990s established that MACHOs do not make an appreciable contribution to the dark matter content of our galaxy.

The tightest constraints on the amount of dark matter in the universe come from cosmological measurements. The frequency and amplitude dependence of the fluctuations in the cosmic microwave background (CMB) measured by WMAP (and in the future by Planck) are sensitive to both the total matter density and the baryon density. The baryon density is also



DARK MATTER SEARCH GOES UNDERGROUND

From a vantage point a half-mile below ground, physicists of the Cryogenic Dark Matter Search have launched a quest to detect the dark matter that pervades the universe. Scientists of CDMS II, an experiment in the Soudan Iron Mine in northeastern Minnesota, hope to discover the weakly interacting massive particles that are leading candidates for the constituents of dark matter. The WIMPs are thought to be particles more massive than protons but interacting	so rarely that thousands would pass through us undetected each second. Only occasionally would a WIMP hit a terrestrial atom, leaving a signal in the CDMS II detector. In the kind of convergence that gets physicists’ attention, the characteristics of WIMPs appear to match those of a particle predicted by supersymmetry, the neutralino. While CDMS II watches for WIMPs, accelerator experiments seek to create neutralinos in particle collisions and measure their properties.
--	---





### Why is there Something Rather than Nothing?

- the darkest question in all philosophy." (William James)
- This problem "can tear the individual's mind asunder," (A. B. Lovell)
- Each of us is "gazed by [this quest]on's] hidden power," (Martin Heidegger)
- "What is it that breathes fire into the equations and makes a universe for them to describe? Why does the universe go to all the bother of existing?" (Stephen Hawking)

### Conclusion: Much Ado about Nothing

The vacuum is a dynamic place:

- Space and time are intertwined.
- Curved by matter, this curvature affects motion.
- A roiling sea of virtual particles created by quantum fluctuations that follow from the uncertainty principle.
- Contains a nonzero energy density that makes up ~70% of the energy of the Universe. The nature of this "dark energy" is almost completely unknown.
- The dark energy propels the accelerating expansion of the universe.

Perhaps the hottest topic in physics right now.

FROM DAVID GERDES' TALK ABOUT NOTHING  
ANN ARBOR, 15 MAR 03

NOTHING IS EVERYTHING

Quantum physics has revealed a stunning truth about “nothing”: even the emptiest vacuum is filled with elementary particles, continually created and destroyed. Particles appear and disappear, flying apart and coming together, in an intricate quantum dance. This far-reaching consequence of quantum mechanics has withstood the most rigorous experimental scrutiny. In fact, these continual fluctuations are at the heart of our quantum understanding of nature.

The dance of quantum particles has special significance today because it contributes to the dark energy that is driving the universe apart. But there's a problem: the vacuum has too much energy. A naive theoretical estimate gives an amount about  $10^{120}$  times too large to fit cosmological observations. The only known way to reduce the energy is to cancel contributions of different particle species against each other, possibly with a new symmetry called supersymmetry. With supersymmetry the result is  $10^{60}$  times better—a huge improvement, but not enough. Even with supersymmetry, what accounts for the other 60 orders of magnitude is still a mystery.

Physics theory predicts that one of the most important particles in the quantum vacuum is the Higgs particle. The Higgs pervades the vacuum, slowing the motion of particles, giving them mass, and preventing atoms from disintegrating. Since it fills the

vacuum, the Higgs itself contributes to the embarrassing factor of  $10^{120}$ .

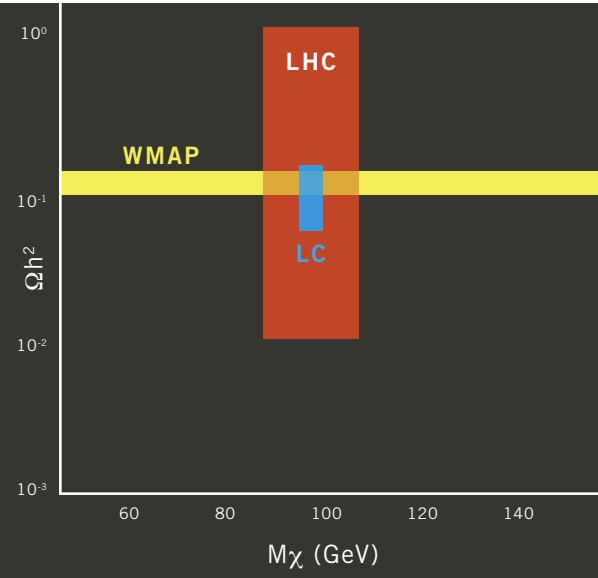
The next accelerators are opening a window on the pivotal role of symmetry in fundamental physics. New discoveries will teach us about the role of the Higgs particle and supersymmetry in defining the vacuum. Such discoveries are key to understanding what tames the quantum vacuum, a topic that is fundamental to any real understanding of the mysterious dark energy that determines the destiny of our cosmos.



MESSIER 31, a thin disk, is the largest galaxy in the neighborhood of the Milky Way. The distribution of mass in M31 and other galaxies led to the recognition that galaxies are filled with dark matter. Credit: Sloan Digital Sky Survey

constrained by the nucleosynthesis models of the early universe. All of these methods suggest that normal baryonic matter can only account for a small fraction, about five percent, of the total matter density.

Scientists are measuring the distribution of dark matter in the universe in a variety of ways: (a) by studying the large-scale distribution of galaxies, as with the Sloan Digital Sky Survey (SDSS); (b) by constraining the dark matter mass power spectrum through weak lensing studies, as by a future Large Synoptic Survey Telescope (LSST) and the Joint Dark Energy Mission (JDEM); and (c) by cataloguing massive clusters of galaxies as a function of redshift, using the Sunyaev-Zeldovitch effect, by the South Polar Telescope and the Atacama Cosmology Telescope.



ACCURACY of WMAP, LHC and linear collider in determining the mass of the neutralino, a dark matter candidate, and so constraining its contribution to the energy density of the universe. Credit: Andreas Birkedal

What is dark matter? Particle physics models suggest that dark matter is either axions (hypothetical new particles associated with QCD), or WIMPs (hypothetical new particles with weak interactions and TeV-scale masses, natural by-products of theories of supersymmetry or extra dimensions). If dark matter particles are relics from the near-total annihilation in the early universe, simple dimensional analysis suggests that the particles originate from physics at the TeV scale. The particle nature of dark matter can be verified by finding the rare events they would produce in a sensitive underground dark matter detector such as CDMS. Such experiments may see products of dark matter particles in our galaxy. Annihilation of TeV-scale dark matter particles might be detected as line radiation in high-energy gamma ray telescopes such as GLAST and VERITAS, or possibly in astrophysical

neutrino detectors such as ICE CUBE. Antiparticles produced in these annihilations might also be detectable by AMS. If dark matter particles are much more massive, they might produce signals in the ultra-high-energy cosmic rays.

However, to understand the true nature of dark matter particles, particle physics experiments must produce them at accelerators and study their quantum properties. Physicists need to discover how they fit into a coherent picture of the universe. Suppose experimenters detect WIMPs streaming through an underground detector. What are they? Are they the lightest supersymmetric particle? The lightest particle moving in extra dimensions? Or are they something else?

Searches for candidate dark matter particles are underway at present-day colliders. If these particles have masses at the TeV scale, they will surely be discovered at the LHC. However, verifying that these new particles are indeed related to dark matter will require the Linear Collider to characterize their properties. The Linear Collider can measure their mass, spin and parity with precision. These results will permit calculation of the present-day cosmic abundance of dark matter and comparison to cosmological observations. If the values agree, it will be a great triumph for both particle physics and cosmology and will extend the understanding of the evolution of the universe back to  $10^{-10}$  seconds after the big bang.



**DARK MATTER MAP** The total mass within giant galaxy cluster CL0025+1654 is the sum of the galaxies themselves, seen in yellow as ordinary luminous matter, plus the cluster's invisible dark matter shown in blue. The cluster's dark matter is not evenly distributed, but follows the clumps of luminous matter closely. *Credit: J.-P. Kneib Observatoire Midi-Pyrenees, Caltech. ESA, NASA*

7  
WHAT ARE NEUTRINOS TELLING US?

The discovery that neutrinos have mass opens a window on physics beyond the Standard Model. The Standard Model cannot accommodate neutrino masses without the introduction of new particles, which themselves raise new questions. In fact, the size of the neutrino masses is consistent with expectations from unified theories that require the new particles for the unification itself.

The most pressing question about neutrinos involves how many different kinds there are. Results from the LSND experiment suggest that there may be more than the canonical three families. If so, that would require a major revision of current understanding. The Mini-

BooNE experiment, now running at Fermilab, will settle this issue by mid-2005.

Even if there are only three kinds of neutrinos, questions remain. What generates the neutrino masses and what are their values? Are neutrinos their own antiparticles? How do different kinds of neutrinos mix? Answering these questions requires precision measurements of the neutrino masses and mixings. Physicists are now studying neutrino mixing at the SNO, KamLAND, K2K and SuperKamiokande experiments. A big step will occur in 2005, when the NuMI/MINOS program at Fermilab begins to probe nu-mu/nu-tau neutrino mixing in a controlled accelerator experiment. In 2006, the CERN-to-Gran Sasso long-baseline neutrino program will begin. The neutrino beam at JPARC is being developed in Japan. In the longer-term future, these experiments and their upgrades, possibly using an off-axis beam, or dedicated reactor neutrino experiments may tell us if a measurement of CP violation in the neutrino sector is feasible. Then researchers might use a neutrino superbeam or neutrino factory to search for it. The detector in such an experiment could also search for proton decay, if located deep underground in a facility such as a National Underground Science and Engineering Laboratory.

Accelerator and reactor oscillation experiments measure mass differences. The masses themselves must be determined by different methods. Neutrinoless double beta decay experiments such as EXO and Majorana can be used to measure the electron neutrino mass to  $\sim 0.01\text{eV}$ , if neutrinos are their own antiparticles. The observation of neutrinoless double beta decay would have far-reaching consequences, raising the possibility that the matter and antimatter could have transformed to each other in the early universe.



The Main Injector Neutrino Oscillation Search experiment, half a mile underground in the historic Soudan iron mine in northern Minnesota, uses a neutrino beam from the Main Injector accelerator at Fermilab to probe the secrets of neutrinos: where do they come from, what are their masses, and how do they transform from one flavor of neutrino into another?	changed into electron neutrinos or tau neutrinos during the split-second trip. More than a trillion man-made neutrinos will pass through the MINOS detector each year. Because neutrinos interact so rarely, only about 1,500 of them will collide with atomic nuclei inside the detector. The rest traverse the detector without a trace. The 200-plus MINOS experimenters use the change, or oscillation, from one type of neutrino to another as the key to investigating neutrino mass.
In the MINOS experiment, a "near" detector on the Fermilab site, outside Chicago, measures the beam of muon neutrinos as it leaves the accelerator. The neutrinos travel 450 miles straight through the earth from Fermilab to Soudan—no tunnel needed. In Minnesota, the 6,000-ton "far" detector searches the beam for muon neutrinos that may have	University scientists from Brazil, France, Greece, Russia, United Kingdom and the United States collaborate on the MINOS experiment using intense neutrino beams from a particle accelerator to discover the secrets of this most elusive of particles.



## THE BIRTH OF THE UNIVERSE

The understanding of the history and future of the universe is deeply entwined with questions of the fundamental nature of matter, energy, space and time. Today the strong connections between the largest and smallest domains are clearly apparent.

### 8

#### HOW DID THE UNIVERSE COME TO BE?

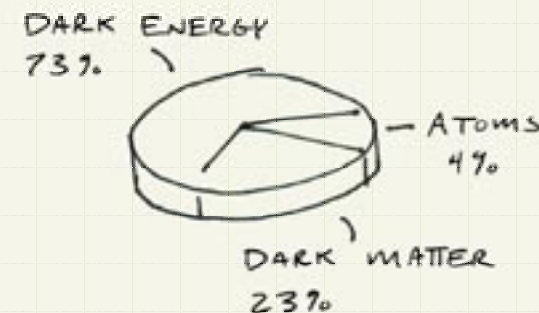
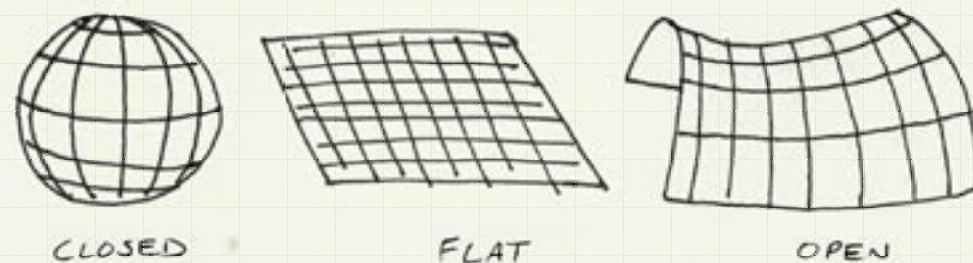
According to current theories of cosmic evolution, the universe begins with an “initial singularity,” a point where all known laws of physics break down. This singularity produced a delicately balanced universe, like a pencil so precisely balanced on its point that it stays upright for 14 billion years. How did the universe reach such a state? How did it get to be so old? Why has it not blasted even further apart, or collapsed back on itself?

For the last two decades, the theory of cosmic inflation has offered a compelling explanation of the start of the big bang. According to this theory, an early phase of accelerated expansion gave rise to the balanced universe we see today. Cosmic inflation is the hand that balanced the pencil on its point. As a by-product, it also produced the seeds that evolved into stars, galaxies, clusters of galaxies and other structures in the universe.

Cosmic inflation presents challenges related to the fundamental questions in this report. One possibility is that cosmic inflation originated with a form of dark energy, akin to the dark energy observed today. If so, what kind of matter produced it? Does this form of matter play a role in unification? How does it relate to extra dimensions? Even more radical is the possibility that space and time changed their nature at the start of the big bang. Does string theory smooth the initial singularity? Which model did nature really choose?

At present, measurements of fluctuations in the cosmic microwave background, especially from WMAP, provide the best evidence in favor of inflation. Constraints on cosmic parameters, such as the curvature of the universe, and the nature of the cosmic structure, are in broad agreement with the predictions of inflationary theory. Eventually, measurements of the polarization of the CMB may permit the detection of the signatures of gravitational

### POSSIBLE GEOMETRY OF UNIVERSE

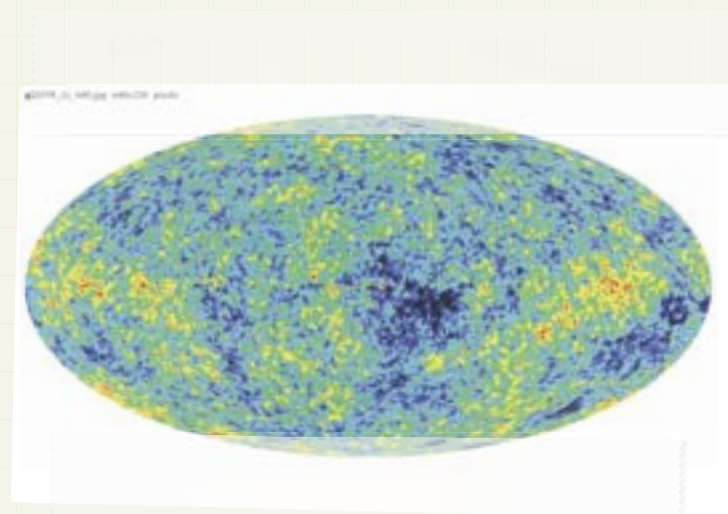


### WMAP

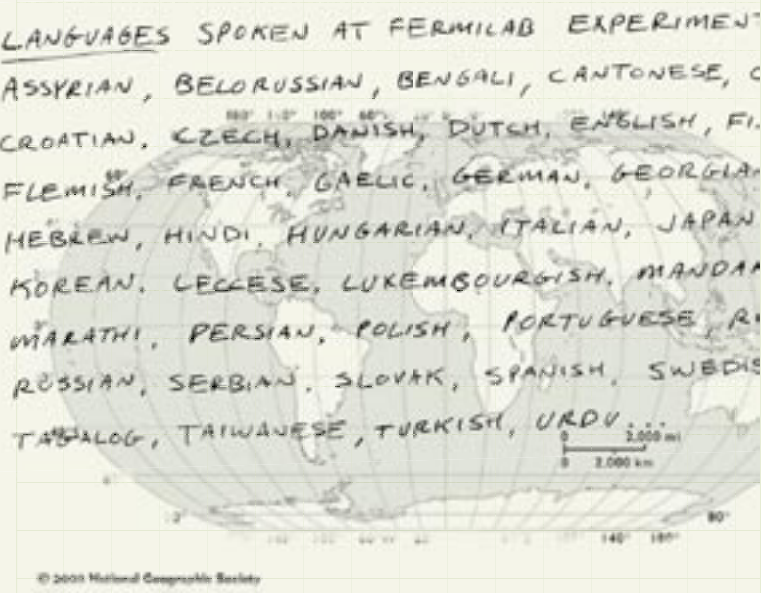
The most extensive measurements of the relic radiation from the early universe are currently being made by the Wilkinson Microwave Anisotropy Probe, a satellite launched in 2001. WMAP consists of an assembly of sensitive microwave receivers to precisely characterize the afterglow of the big bang. Observations from WMAP have confirmed that the space-time geometry of the universe is flat, like a sheet of paper, and have then yielded the age of the universe—about 13.7 billion years—with a precision of about one percent. Further investigations of the variation of the microwave background radiation at small angular scales have shown that normal visible matter plus cold dark matter can account for only about 30 percent of the content of our

universe. The remaining 70 percent is dark energy. WMAP has accurately determined the content of normal visible matter at only four percent, in remarkable agreement with the value determined using data from the much earlier era of nucleosynthesis. The whole revolutionary picture of our universe, including dark matter and dark energy, matches beautifully with the WMAP data.

The flight hardware and software for WMAP are produced in a partnership between NASA's Goddard Space Flight Center and Princeton University, under the scientific supervision of a team whose institutions also include UCLA, the University of Chicago, the University of British Columbia, and Brown University.



WMAP: DETAILED FULL-SKY MAP  
OF THE OLDEST LIGHT IN THE UNIVERSE!



AN INTERNATIONAL SCIENCE

The scientific opportunities provided by particle physics experiments bring hundreds of scientists from every corner of the globe to work together at experiments all over the world. Both the technical scale and the cost of today's accelerators and experiments put them beyond the reach of any single nation's ability to build or operate. Particle physics	projects now take shape as international endeavors from their inception. These scientific collaborations take on new significance as beacons for free and open exchange among men and women of science from all nations. They offer an inspiring model for cooperation from a field known for its leadership in international collaboration.
--	--



waves produced during the epoch of inflation, which could provide information about the nature of the scalar field that produced inflation.

After the big bang, the universe expanded and cooled to reach its present state. Along the way, the universe passed through a series of phase transitions in which various particles froze out, as water turns to ice as it cools. These phase transitions drove some of the most important epochs of cosmic history. For example, a phase transition may be what drove cosmic inflation. Phase transitions might produce “cosmic defects,” such as strings and texture and other exotic forms of matter, that could explain ultra-high-energy cosmic rays, dark matter and perhaps even dark energy.

The LHC will illuminate the electroweak phase transition, where most of the known particles acquired their masses. Better understanding of this phase transition will allow scientists to push closer to the big bang itself. Indeed, it is likely that the electroweak phase transition is the ultimate source of the matter-antimatter asymmetry we see in the universe today. Discoveries of new particles and new interactions will illuminate this story and determine if it is correct. Moreover, the account of cosmic evolution must incorporate any discoveries of new symmetries or new dimensions.

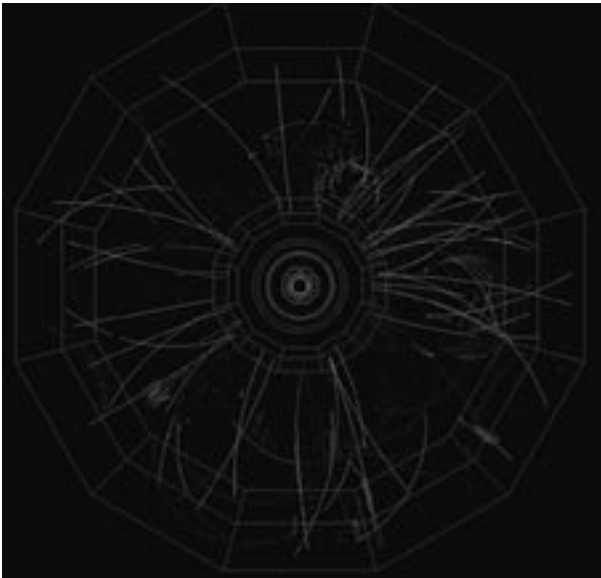
Currently, the most intensely studied cosmic phase transition is connected with quantum chromodynamics (QCD), the theory of the nuclear force. During the QCD phase transition, the baryonic matter in the present universe condensed from a plasma-like state of quarks and gluons. The Relativistic Heavy Ion Collider (RHIC) facility at BNL is currently creating



**ATLAS** Construction of the ATLAS detector at CERN's LHC.  
*Credit: CERN*

collisions of heavy ions to study quark-gluon plasma; the laboratory plans upgrades to enhance these studies. The Lattice Computational Facilities will enable calculations furthering the understanding of the RHIC data and the conditions during this epoch in the evolution of the early universe.

The synthesis of all the elements in the world involves nuclear reactions. Given the temperatures and particle densities in stellar objects and in cataclysmic stellar explosions, these reactions often occur in unstable nuclei. The Rare Isotope Accelerator, RIA, will provide the tool for the terrestrial study of the nuclear reactions that drive these events and will help us understand how and where in the universe nature synthesizes the elements.

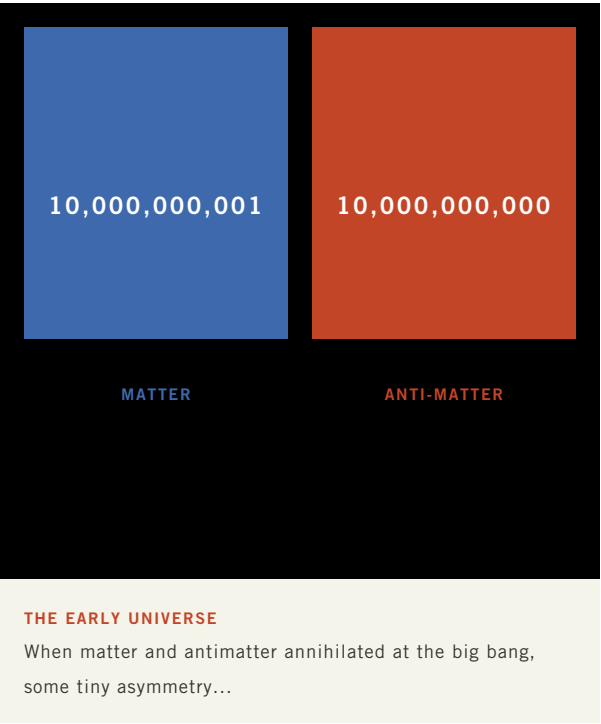


**RHIC** An end view of a collision between deuterons and gold ions captured by the STAR detector at Brookhaven's Relativistic Heavy Ion Collider. *Credit: Brookhaven*

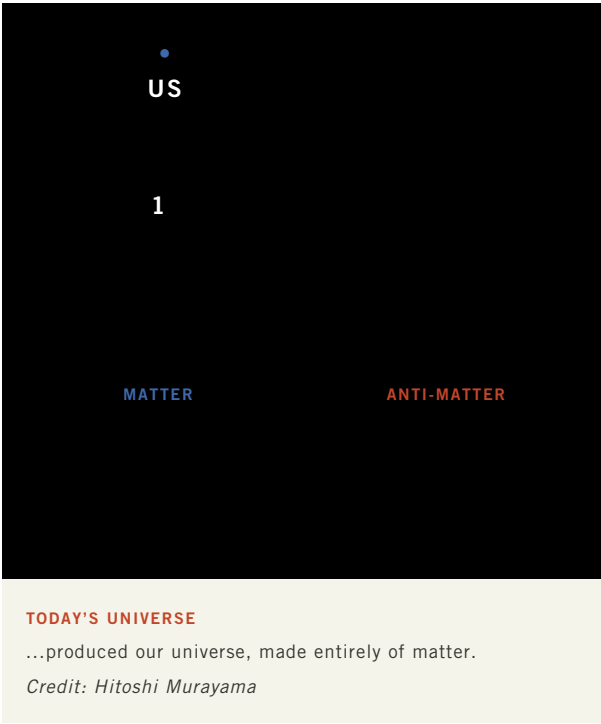
9  
WHAT HAPPENED TO THE ANTIMATTER?

A fundamental question in the evolution of the universe is What happened to the antimatter? Matter and antimatter were almost certainly produced equally at the birth of the universe. However, as the soup of the hot early universe cooled, equal amounts of matter and antimatter would have combined and annihilated. Instead, an excess of matter has remained to form the galaxies and stars that constitute the universe. To preferentially remove antimatter in the universe, CP symmetry must be violated, causing antimatter to behave slightly differently from matter. Experimenters discovered CP violation in the neutral K mesons in 1964, and in the B mesons in 2001.

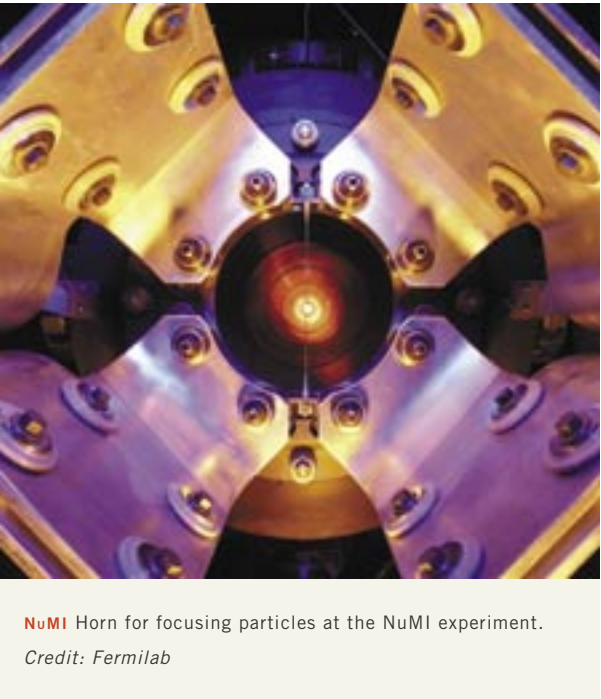




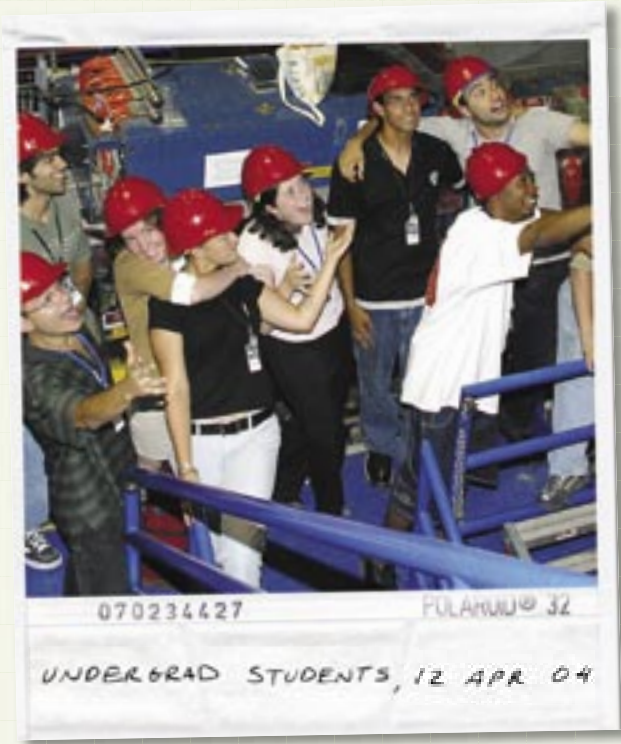
The Standard Model can accommodate the phenomenology of CP violation in quarks, because there are at least three generations of quarks; and because there is mixing between the quark flavors when they interact via the weak interaction. The CP violation measured in the B mesons at BaBar and Belle, along with a wealth of studies on quark flavor mixing over the past 20 years, are all consistent with this phenomenology. However current knowledge of CP violation is incomplete and insufficient by many orders of magnitude to account for the primordial matter-antimatter asymmetry of the universe. Present and planned accelerator experiments are aimed at discovering other sources of CP violation that make matter and antimatter behave differently. It may appear in quarks-or in neutrinos. Its source may lie in the properties of the Higgs boson, in supersymmetry or even in extra dimensions.



CP violation in reactions that change the flavor of quarks is being measured with strange quarks (K meson decays) and bottom quarks (B meson decays). Ongoing and planned experiments include KOP10 at BNL (K decay); BaBar at SLAC, Belle at KEK ( $B_d$  decay), BTeV at Fermilab and LHC-b at CERN ( $B_d$  and  $B_s$  decay). Pinning down precisely the role of CP violation in the quarks is a critical step in solving the puzzle of the fate of primordial antimatter. Experiments have so far shown that, by itself, CP violation in quarks from flavor mixing in the Standard Model is probably not the sole source of the matter-antimatter asymmetry observed in the universe. Current and future B physics experiments will be sensitive to sources of CP violation beyond the Standard Model.



The discovery that neutrinos have mass opens up the search for CP violation in lepton reactions. Neutrino mass can, in principle, turn matter into antimatter and back, and can change the balance between them. Experiments are required to discover the role of neutrinos in the antimatter question. The MINOS experiment at Fermilab and reactor-based neutrino oscillation experiments will measure the parameters of neutrino oscillation. If the oscillation parameters are favorable, a neutrino superbeam facility with a large underground experiment will detect CP violation in neutrinos. Such a large detector, if sufficiently far underground, for example at a potential Deep Underground Science and Engineering Laboratory, could also serve as a next-generation proton-decay experiment



EDUCATION AND OUTREACH

Particle physics research is a journey of exploration into the mystery and beauty of the universe. A universe made up of tiny vibrating strings and extra dimensions, filled with dark matter and dark energy, could soon become our new reality. Excitement and fascination have always attracted young people to physics and science, and the questions asked in modern particle physics are glamorous, provoking wonder and dream. Linking today's high school students will be the researchers on many of the projects described in this report. By connecting science teachers with the excitement	of research, we can help them pass on their enthusiasm to their students. Among such opportunities, QuarkNet, for example, brings high school students and teachers to the frontier of 21st century research by involving them directly in the programs of the major particle physics laboratories. Through university scientists across the country, QuarkNet is connected to high-energy physics experiments operating at Fermilab, SLAC, CERN and other research centers. Working with the educational links in the laboratories and universities, the nation's science teachers can catalyze the interest of our young people in today's revolutionary science of particle physics.
---	---

# CHAPTER IV CONCLUSIONS

WHAT IS THE NATURE OF THE UNIVERSE  
AND WHAT IS IT MADE OF?

WHAT ARE MATTER, ENERGY,  
SPACE AND TIME?

HOW DID WE GET HERE AND  
WHERE ARE WE GOING?

Particle physics is in the midst of a great revolution. Modern data and ideas have challenged long-held beliefs about matter, energy, space and time. Observations have confirmed that 95 percent of the universe is made of dark energy and dark matter unlike any we have seen or touched in our most advanced experiments. Theorists have found a way to reconcile gravity with quantum physics, but at the price of postulating extra dimensions beyond the familiar four dimensions of space and time.

As the magnitude of the current revolution becomes apparent, the science of particle physics has a clear path forward. The new data and ideas have not only challenged the old ways of thinking, they have also pointed to the steps required to make progress. Many advances are within reach of our current program; others are close at hand. We are extraordinarily fortunate to live in a time when the great questions are yielding a whole new level of understanding. We should seize the moment and embrace the challenges.



Photo Credit: Fermilab

CHAPTER V  
SUMMARY TABLES

The following two tables summarize selected facilities of the U.S. program whose primary physics goals align most directly with the report’s nine questions. Other important facilities are included in the text.

LEGEND: THE QUESTIONS

- 1
- Are there undiscovered principles of nature :  
new symmetries, new physical laws?
- 2
- How can we solve the mystery of dark energy?
- 3
- Are there extra dimensions of space?
- 4
- Do all the forces become one?
- 5
- Why are there so many kinds of particles?
- 6
- What is dark matter?  
How can we make it in the laboratory?
- 7
- What are neutrinos telling us?
- 8
- How did the universe come to be?
- 9
- What happened to the antimatter?

PRIMARY US PHYSICS  
PROGRAM OF MAJOR  
FACILITIES

This table summarizes the physics goals of the major facilities of the US program whose primary physics goals align most directly with the report's nine questions.

		Tevatron	LHC	Linear Collider	NUMI / MINOS	$\nu$ Superbeams	BaBar	BTeV	JDEM	RHIC	Proton Decay
Undiscovered Principles?	1	X	X	X			X	X			
Dark Energy?	2		X	X					X		
Extra Dimensions?	3		X	X							
Unified Forces?	4			X							X
Why So Many Particles?	5	X					X	X			
Dark Matter?	6			X					X		
Neutrinos?	7				X	X					
Origin of Universe?	8		X							X	
Antimatter?	9					X	X	X			

PRIMARY US PHYSICS  
PROGRAM OF  
SELECTED SMALLER  
FACILITIES

This table summarizes the physics goals of selected smaller facilities of the US program whose primary physics goals align most directly with the report's nine questions.

		Mini-BooNE	MECO	Reactor $\nu$ Experiments	CLEO-c	KOP10	Neutrinoless Dbl Beta Decay	SDSS	LSST	Undgrnd Dark Matter Expts	WMAP	CMB Polarization	Lattice Computational Facilities	Precision Gravity
Undiscovered Principles?	1	X												
Dark Energy?	2								X		X			
Extra Dimensions?	3												X	
Unified Forces?	4						X							
Why So Many Particles?	5		X		X								X	
Dark Matter?	6							X	X	X	X			
Neutrinos?	7	X		X			X							
Origin of Universe?	8										X	X	X	
Antimatter?	9					X								

CHARGE TO THE COMMITTEE

Professor Frederick Gilman  
Chair, HEPAP  
Carnegie-Mellon University  
5000 Forbes Avenue  
Pittsburgh, PA 15213



October 22, 2003

Dear Professor Gilman:

Recent scientific discoveries at the energy frontier and in the far reaches of the universe have redefined the scientific landscape for cosmology, astrophysics and high energy physics, and revealed new and compelling mysteries. We are writing to ask the High Energy Physics Advisory Panel (HEPAP) to take the lead in producing a report which will illuminate the issues, and provide the funding and science policy agencies with a clear picture of the connected, complementary experimental approaches to the truly exciting scientific questions of this century. The report should elucidate how the questions being asked in particle physics overlap with those being asked by other communities. Further, we are particularly interested in the role that accelerators will play in addressing the important questions and the complementary roles played by other experimental techniques.

We request that HEPAP form a committee that will write a report identifying and addressing the key questions now faced by high energy physics, particle astrophysics and cosmology:

- What are the general methods and technologies that can answer these questions, and what is the particular contribution made by particle accelerators?
- What is the current status of scientific efforts in these areas and what are the near-term prospects for advances?
- Explain the connections between various approaches to this research. How can the results from one type of investigation impact the science of another experiment ? For example, discuss the interrelation of searches for dark matter .

The membership of the committee should be drawn broadly from the communities in particle physics, nuclear physics, cosmology, astrophysics and related fields that are actively involved in this science and can give independent advice on the relative strengths of the various approaches considered. We would like a brief report which encompasses the most important scientific questions and addresses the issues outlined above in a summary fashion. We recognize that, given the complexity of interconnections between fields, further studies may be needed to give a more complete picture of these evolving areas. We appreciate your advice on the appropriate “next steps” to follow-up on your report.

We look forward to the creation of this committee in the near future. We would like to have a status report on the work of the committee by the end of 2003, with a final report to HEPAP early in 2004.

We wish you success in this challenging and important endeavor.

Sincerely,

Raymond L. Orbach  
Director  
Office of Science  
U.S. Department of Energy

Michael Turner  
Assistant Director  
Mathematical and Physical Sciences  
National Science Foundation

cc: Ben Weakley, SC-4  
Robin Staffin, SC-20  
Bruce Strauss, SC-20  
Joseph Dehmer, NSF



FACILITIES GLOSSARY

**ACT** Atacama Cosmology Telescope. This is a proposed 6 meter telescope that will look at millimeter radiation from the early universe to map the formation of cosmic structure. [www.hep.upenn.edu/act](http://www.hep.upenn.edu/act)

**AMS** Alpha Magnetic Spectrometer is a particle detector being constructed for the International Space Station. It will observe properties of electrons, positrons, protons, antiprotons and nuclei in high-energy radiation in space. [cyclo.mit.edu/ams](http://cyclo.mit.edu/ams)

**BABAR** This experiment, which focuses on B mesons, uses the PEP II  $e^+e^-$  storage ring at the Stanford Linear Accelerator Center (SLAC). One of its primary goals is to investigate why the universe appears to contain matter but no antimatter. BaBar and Belle discovered the first evidence of CP violation outside of the K meson system. [www.slac.stanford.edu/BFR00T](http://www.slac.stanford.edu/BFR00T)

**BELLE** This experiment, which focuses on B mesons produced at the KEKB  $e^+e^-$  collider, is located at the KEK laboratory in Japan. One of its primary goals is to investigate why the universe appears to contain matter but no antimatter. BaBar and Belle discovered the first evidence of CP violation outside of the K meson system. [bsunsrv1.kek.jp](http://bsunsrv1.kek.jp)

**BTEV** A proposed experiment at the Fermilab Tevatron to study decays of B mesons in detail in hadron collisions. BTeV and LHC-b will do unique

studies of the  $B_s$  mesons and explore their role in the matter antimatter asymmetry of the universe. [www-btev.fnal.gov](http://www-btev.fnal.gov)

**CNGS** CERN Neutrinos to Gran Sasso. The next phase of the CERN neutrino program uses a neutrino beam from CERN to detectors in the Gran Sasso tunnel to study neutrino oscillations. [proj-cngs.web.cern.ch](http://proj-cngs.web.cern.ch)

**CLEO-C** This experiment is based at the CESR accelerator. It is focused on high-precision studies of charm mesons to validate precision lattice calculations. [www.lepp.cornell.edu](http://www.lepp.cornell.edu)

**CMB POLARIZATION** A variety of experiments at various stages of development are studying the polarization of the cosmic microwave background. Measurements of partial polarization of the CMB on small scales will yield fundamental information about the evolution of the early universe and provide key tests of our current theories.

**GLAST** Gamma Ray Large Area Space Telescope. A satellite-based experiment, presently under construction, which will detect high-energy gamma-rays from cosmic sources. GLAST measurements will provide crucial data on acceleration mechanisms in astrophysical settings, and could provide constraints on dark matter models through searches for annihilation features at gamma-ray energies. [www-glast.stanford.edu](http://www-glast.stanford.edu)

**ICECUBE** IceCube is a one-cubic-kilometer international high-energy neutrino observatory being built and installed in the ice below the South Pole Station. IceCube will search for neutrinos from gamma ray bursts, black holes and neutron stars. IceCube will also provide important constraints on models of dark matter. [icecube.wisc.edu](http://icecube.wisc.edu)

**JDEM** Joint Dark Energy Mission. This is a proposed satellite-based experiment for finding and studying supernovae to determine the evolution of the dark energy in the universe. JDEM will also use weak and strong lensing of distant galaxies to better measure both the dark energy and dark matter distributions in the universe.

**JPARC** Japan Proton Accelerator Research Complex. A high-intensity proton accelerator facility in Japan that will produce intense neutrino beams to study neutrino oscillation parameters. [jkj.tokai.jaeri.go.jp](http://jkj.tokai.jaeri.go.jp)

**KAMLAND** A 1000-ton scintillation detector located in the Kamioka mine in Japan. KamLAND has observed oscillations of antineutrinos from nearby nuclear reactors. [www.awa.tohoku.ac.jp/html/KamLAND/](http://www.awa.tohoku.ac.jp/html/KamLAND/)

**K2K** This experiment aims a neutrino beam from the KEK accelerator to the SuperKamiokande detector. It has given evidence for neutrino oscillations consistent with those seen in atmospheric neutrinos. [neutrino.kek.jp](http://neutrino.kek.jp)

**KOPIO** An approved experiment that searches for a rare decay of neutral K-mesons into a neutral pion and two neutrinos. This decay provides valuable information about quark mixing and CP violation. [www.bnl.gov/rsvp/](http://www.bnl.gov/rsvp/)

**LATTICE COMPUTATIONAL FACILITIES** Proposed computational facilities used to calculate strong interaction dynamics.

**LHC LARGE HADRON COLLIDER** A 14 TeV proton-proton collider under construction at the CERN laboratory in Geneva, Switzerland. Two large international detectors, ATLAS and CMS, will explore new physics at the TeV scale. The LHC opens a new high-energy frontier for the field. [lhcb.web.cern.ch/lhc/](http://lhcb.web.cern.ch/lhc/)

**LHC-B** An experiment under construction at CERN to study B mesons in detail in hadron collisions. BTeV and LHC-b will do unique studies of the  $B_s$  mesons and explore their role in the matter antimatter asymmetry of the Universe. [lhcb.web.cern.ch/lhcb/](http://lhcb.web.cern.ch/lhcb/)

**LIGO** Laser Interferometer Gravitational Wave Observatory. A facility to detect astrophysical gravitational waves consisting of two widely separated installations in Hanford, Washington and Livingston, Louisiana, operated in unison as a single observatory. [www.ligo.caltech.edu](http://www.ligo.caltech.edu)

**LINEAR COLLIDER** A proposed international TeV  $e^+e^-$  accelerator based on two linear accelerators whose beams collide head-on. The linear collider will make precision measurements of the new TeV-scale physics discovered by LHC. [blueox.uoregon.edu/~lc/wwstudy/](http://blueox.uoregon.edu/~lc/wwstudy/)

**LSST** Large-aperture Synoptic Survey Telescope. A proposed wide-field ground-based optical telescope that will survey a major fraction of the sky every few days. It will enable studies of weak gravitational lensing that will provide constraints on dark matter and dark energy in the universe. Other applications include the search for near-Earth asteroids that could potentially collide with the Earth. [www.lsst.org](http://www.lsst.org)

**MECO** Muon to Electron Conversion Experiment. An approved experiment to search for the conversion of muons to electrons in the presence of nuclei. This would be a signal of lepton flavor violation. [www.bnl.gov/rsvp/](http://www.bnl.gov/rsvp/)

**MINIBOONE** Mini-Booster Neutrino Experiment. This experiment, based at Fermilab, is searching for oscillations between muon and electron neutrinos using a beam of neutrinos. [www-boone.fnal.gov](http://www-boone.fnal.gov)

**NEUTRINOLESS DOUBLE BETA DECAY** Sensitive experiments to search for evidence of a lepton number violating decay that would indicate the neutrino is its own antiparticle.

**NUMI/MINOS** Neutrinos at the Main Injector/ Main Injector Neutrino Oscillation Search. An experiment coming on line in 2005 that shoots neutrinos from the Main Injector at Fermilab toward an underground target 730 km (430 miles) away in Soudan, Minnesota. [www-numi.fnal.gov](http://www-numi.fnal.gov)

**PLANCK** A European space mission under construction to make precision measurements of fluctuations in the cosmic microwave background. [www.rssd.esa.int/index.php?project=PLANCK](http://www.rssd.esa.int/index.php?project=PLANCK)

**PRECISION GRAVITY** Small scale experiments that test the theory of gravity to sub-millimeter scales.

**PROTON DECAY EXPERIMENTS** Proposed very large underground experiments that are sensitive to the decay of the proton. These detectors can also be used as detectors for accelerator-based neutrino oscillation experiments.

**REACTOR  $\bar{\nu}$  EXPERIMENTS** Proposed experiments to use neutrinos from reactors to measure parameters of neutrino oscillation.

**RIA** Rare Isotope Accelerator. A proposed facility to produce and explore new rare isotopes that cannot be found on Earth. [www.orau.org/ria](http://www.orau.org/ria)

**RHIC** Relativistic Heavy Ion Collider. A Brookhaven National Laboratory accelerator that collides heavy ions with 100 GeV per nucleon per beam to create new states of matter. [www.bnl.gov/RHIC](http://www.bnl.gov/RHIC)

**SDSS** Sloan Digital Sky Survey. A ground-based telescope that maps the sky, determining positions and absolute brightness of over 100 million celestial objects. [www.sdss.org](http://www.sdss.org)

**SNO** Sudbury Neutrino Observatory. A heavy-water Cherenkov detector located 2 km below the surface in a nickel mine in Ontario. SNO has given compelling evidence for oscillation of electron neutrinos to other flavors, providing the explanation for the solar neutrino problem. [www.sno.phy.queensu.ca](http://www.sno.phy.queensu.ca)

**SPT** South Pole Telescope. An 8 meter precision sub-millimeter-wave telescope that is under construction. SPT will do a large scale survey of distant galaxy clusters and map the evolution of large-scale structure in the very early universe. [astro.uchicago.edu/spt/](http://astro.uchicago.edu/spt/)

**SUPERKAMIOKANDE** This experiment, located in the Kamioka mine in Japan, is the successor to the smaller Kamiokande experiment. SuperKamiokande provided compelling evidence for neutrino oscillations in neutrinos created by cosmic rays bombarding the Earth's atmosphere. [www-sk.icrr.u-tokyo.ac.jp/doc/sk](http://www-sk.icrr.u-tokyo.ac.jp/doc/sk)

**TEVATRON** The 2 TeV proton-antiproton collider at Fermilab, currently the world's highest-energy accelerator. At the Tevatron, two large general purpose detectors, CDF ([www-cdf.fnal.gov](http://www-cdf.fnal.gov)) and D0 ([www-d0.fnal.gov](http://www-d0.fnal.gov)), are studying the particles of the standard model and looking for new physics beyond it. CDF and D0 discovered the top quark.

**UNDERGROUND DARK MATTER DETECTORS** Sensitive underground detectors that search for dark matter through its interactions with matter.

**VERITAS** Very Energetic Radiation Imaging Telescope Array System. A ground-based gamma-ray observatory that detects optical flashes in the atmosphere produced by air showers generated by cosmic gamma-rays in the GeV—TeV energy range. VERITAS studies cosmic acceleration mechanisms and exotic objects such as supernovae remnants and active galactic nuclei. It will also search for dark matter in our galaxy. [veritas.sao.arizona.edu](http://veritas.sao.arizona.edu)

**$\nu$  SUPERBEAMS** Possible future accelerators capable of producing intense beams of neutrinos to make precision measurements of neutrinos and perhaps be sensitive to CP violation in the neutrinos.

**WMAP** Wilkinson Microwave Anisotropy Probe. A satellite mission that has made the most sensitive measurements to date of the fluctuations in the cosmic microwave background, leading to tight constraints on cosmological parameters. [map.gsfc.nasa.gov](http://map.gsfc.nasa.gov)

Quantum Universe was prepared in response to a charge to the High Energy Physics Advisory Panel by Dr. Raymond L. Orbach, Director of the Office of Science of U.S. Department of Energy and Dr. Michael Turner, Assistant Director for Mathematics and Physical Science of the National Science Foundation.

**HEPAP**  
The High Energy Physics Advisory Panel has advised the federal government on the national program in experimental and theoretical high-energy physics research since 1967. HEPAP reports directly to the director of the U.S. Department of Energy's Office of Science and to the assistant director, Mathematical and Physical Sciences, of the National Science Foundation.



**Office of Science**  
U.S. Department of Energy

**THE OFFICE OF SCIENCE**  
The Office of Science of the U.S. Department of Energy is the single largest supporter of basic research in the physical sciences in the United States, providing more than 40 percent of total funding for this vital area of national

importance. It oversees—and is the principal federal funding agency for—the nation's research programs in high-energy physics, nuclear physics and fusion energy sciences.

The Office of Science manages fundamental research programs in basic energy sciences, biological and environmental sciences and computational science. In addition, the Office of Science is the federal government's largest single funder of materials and chemical sciences, and it supports unique and vital parts of U.S. research in climate change, geophysics, genomics, life sciences, and science education. The Office of Science continues to strengthen partnerships with other science agencies, including NSF and NASA, in support of basic research in the physical sciences.



**THE NATIONAL SCIENCE FOUNDATION**  
Mathematical and Physical Sciences Directorate  
The Mathematical and Physical Sciences Directorate (MPS) supports a strong and diverse portfolio of research and education in mathematics, astronomical science, physics, chemistry and materials research. The purpose of this work is both to deepen our understanding of the physical universe and to use this understanding in service to society. MPS is involved in many long-standing, very valuable partnerships with its sister US agencies, including DOE, NASA and NIH, as well as with international consortia. This mode of operation is sure to be augmented in the coming years.



**NASA**  
The National Aeronautics and Space Administration is the federal agency responsible for the space program and long-term aerospace research. Founded in 1958, NASA is a civilian organization that conducts or oversees research into both civilian and military aerospace systems.



What is  $W + \gamma\gamma$  rate?

$200 \text{ pb}^{-1}$ ,  $\text{Accep} = 0.5$  (guess)

$\text{BR}(W \rightarrow e) = 0.1 \rightarrow 1 \text{ event}$   
 $= 100 \text{ fb} ?$